ASTR 1020: Stars & Galaxies April 23, 2008

- Reading: Chapter 23, sections 23.3-23.4 .
- *MasteringAstronomy* Homework on The Fate of the Universe is due April 30th.
- 2 more extra credit opportunities at Fiske Planetarium this week and next!
- Meet at Fiske Planetarium next Monday.

<section-header> Astronomy Picture of the Day Image: Strain Strai

Last Time

The Creation of the Universe

- 1.) Planck Era: before physics as we know it existed
- 2.) GUT & Electroweak Eras: Four forces came into being
- 3.) Particle Era: Origin of matter
- 4.) Nucleosynthesis (fusion) Era: Helium is born

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Today

- From 3 minutes after the Big Bang Until Now
- Cosmic Microwave Background
- Evidence for the Big Bang Theory

Era of Nuclei

- Dense clouds of protons (hydrogen nuclei), helium nuclei, electrons, neutrinos, photons
- Temperatures too hot for electrons to combine with protons– electrons ionized by energetic photons
- Universe is made of naked nuclei, not atoms with nuclei + electrons
- · Lasts for about 380,000 years

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The Cosmic Microwave Background

- Discovered by accident by Arno Penzias and Robert Wilson in 1965
- Direct evidence for HOT early universe
 → Big Bang
- 1978 Nobel Prize

Clicker Question: Remembering Wien's Law • Wavelength maximum = 2,900,000 nm / T (K)

- If T = 2900 K, what is wavelength of the peak of the thermal spectrum?
- a) (pink) 100 nm
- b) (yellow) 1000 nm
- c) (green) 1,000,000 nm



Remember that visible light is 400-700nm

- → Near infrared, spilling towards red/visible, similar to the coolest red stars.
- We should see a red glow all around us?



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- 1+z ~ 1000
- Wavelength will be BIGGER because universe is now bigger:
- Wavelength emitted x (1+z) = wavelength observed now
- 1,000 nm x 1000 = 1,000,000 nm = 1 mm Microwaves!





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