### ASTR 1020: Stars & Galaxies March 19, 2008

- Reading: Chapter 20, sections 20.3.
- Meet Friday at Fiske Planetarium Hubble's Expanding Universe (bring your clicker)





# A Universe Full of Galaxies

- · Galaxies are classified into basic types
- Use both shapes and star properties
- Sizes range from giants, through biggish (like the Milky Way), through dwarfs





# Elliptical ~15% of galaxies

- Round or slightly flattened
- Very little cold gas (no 21-cm emission), dust, or young stars
- Reddish/yellow color = old stars (red giants, red main sequence)







# Where they live

 Spirals– mostly in groups (3-10 galaxies)







## Mapping the Universe: We need Distances to Galaxies!

So far- Parallax

New methods: standard candles

- 1.) Make some measure of an object which identifies its luminosity
- 2.) Use this luminosity and measure apparent brightness to infer distance to it

# <section-header> **1.) Main sequence fitting** Start with a cluster distance known via parallax (upper) Compare with other clusters (lower) Which is more distant- the upper or lower?







- Clicker Question: Two Cepheid stars, Fred and Barney, have the same apparent brightness. Fred has a period of 5 days, and Barney of 10 days. Which is closer?
- a) Fred
- b) Barney

### • A) Fred

- Fred has a shorter period and so must be less luminous (hummingbird)
- Less luminous but the same apparent brightness means that Fred is closer to us

# **Cepheids as Standard Candles**

- Measure period of variability
- From periodluminosity relation, infer the luminosity
- Compare with apparent brightness and determine distance

