





~Advanced Camera for Surveys (ACS) ~Wide Field Planetary Camera 2 (WFPC2) ~Space Telescope Imaging Spectrograph (STIS)

Multi-Object Spectrometer (NICMOS)

~Corrective Optics Space Telescope Axial Replacement (COSTAR)





~Wavelengths: 1200-10,000 Angstroms (ultraviolet to near infrared)

- 2)High Resolution Camera













Carina Nebula





~Records images through a selection of 48 color filters

~Spectral range varies from far-ultraviolet, to visible, to near-infrared

~Most used instrument in first 13

~Built-in corrective optics

STATS:

results!!!....



Distant Galaxy Collision triggering massive amounts of star formation







Space Telescope Imaging Spectrograph



- STATS:
- Field of View: MAMA 25*25 arcsec; CCD 50*50 arcsec Wavelength Range: 1150-10,000 Angstroms
- Installed in 1997 to replace the GHRS (Goddard High Resolution Spectrograph)
- Chemical compositions and abundances, temperatures, velocities (radial and rotational), magnetic fields
- ~2 Modes:
- 1) long slit spectroscopy
- 2) echelle spectroscopy







NICMOS

~three adjacent cameras (operate independently)

~infrared wavelengths only- operates under 77 degrees Kelvin

~detectors cooled



Dimensions: 2.2m * 0.89m * 0.89m Wavelength: 8000-25,000 Angstron Med: 17.5 * 17.5 arcseconds







^ NICMOS image of the centre of the Milky Way galaxy < NICMOS image of

the Orion Nebula



< 4 Galaxy Collision

Young massive star with 6 "baby stars" within a tenth of a light year





NICMOS Cooling System (NCS)



~experimental technology

~circulates cold Neon gas through NICMOS cryostat _____

~low vibration technology

~more stable and optimal temperatures than N ice





In celebration of the 100 Billionth revolution of the cooler turbine...



HUDDIe Space PRC97.11 - ST Sci OPO - May 12, 1997 R. Thompson (Univ. Arizona), D. Hines (Univ. Arizona), R. Sahai (JPL) and NASA

Corrective Optics Space Telescope Axial Replacement





~installed in 1993 to correct the aberration of HST's primary mirror

~instrument is now performing just about up to its theoretical optimum (resolution and imaging)

~since Hubble's servicing mission in 2002, COSTAR's mission has been complete and COSTAR will be removed from the Hubble during the 2008 servicing mission

Hubble Servicing Mission 4



>I got a picture of this cat by searching HST, COS and WFPC 3.... >.< ...

<u>New Instruments:</u>





Wide Field Planetary Camera 3

- ~planned installation during Hubble servicing mission #4 in 2008
- ~two channels will allow imaging from near infrared
- to near ultraviolet
- ~Much better discovery efficiency in both Infrared and Ultraviolet wavelengths



Discovery efficiency =field view * optical throughput $\sim\!mostly$ being constructed at Goddard Space flight Centre and Ball Aerospace

Cosmic Origins Spectrograph



- ~20 times more sensitive in far ultraviolet
- ~Science Objectives (core issues of NASA's Origins program)
- •Origin of large scale structure and intergalactic medium

•Formation, evolution, ages of galaxies

•Stellar and planetary origins and cold interstellar medium _____

Age/Expansion of the Universe

- ~Measuring the apparent brightness of distant supernovae – Accelerating Universe
- ~Test against oldest star ages

















Hubble Space Telescope



~Groundbreaki ng Observations and Discoveries

~New technologies that will be used in new telescopes

~Lots of work on HST has been done locally at Ball Aerospace!