

ASTR 1020: Stars & Galaxies November 18, 2013

- Reading: Chapter 23, Sections 23.3.
- *Mastering Astronomy* Homework on *Galaxy Evolution* is due Nov. 15th.
- **Exam 3 is on Wednesday!** Review session Tuesday night at 7 pm in EDUC 155.
- Extra credit telescope observing at SBO. Thursday night at 7 pm.



MAVEN to Mars!



Exam 3 will cover

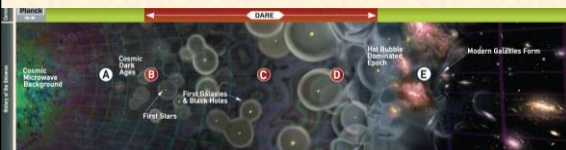
- All material discussed in class, readings, recitations, and homeworks up through today's class.
- Textbook: Chapters 19, 20, 21, 23.
- *Mastering Astronomy* Homeworks on *The Milky Way, Galaxies and Galaxy Evolution*.

The Day of the Exam

- Bring a #2 pencil and eraser or pen.
- One sheet of paper with your notes and study hints.
- Bring a calculator if you think you'll need one.
- Please be prepared to get started right away at 1 pm!

Today's Lecture: The Dark Ages & Cosmic Dawn

- The first stars & galaxies in the Universe.
- The *Dark Ages Radio Explorer* – new space mission led by CU.
- Exploring the Universe from the Farside of the Moon.



Clicker Question

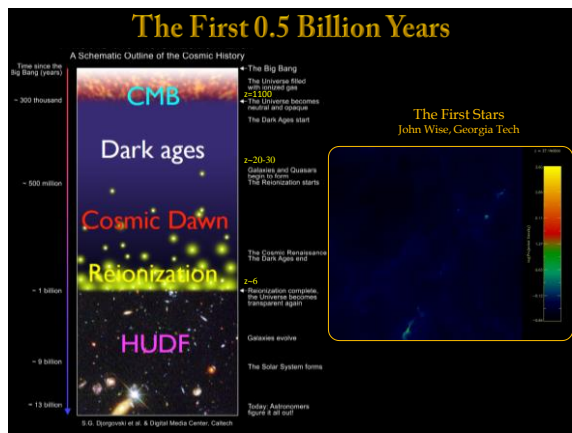
What kind of measurement does not tell us the mass of a cluster of galaxies?

- Measure velocities of cluster galaxies
- Measure total mass of cluster's stars
- Measure temperature of its hot gas
- Measure distorted images of background galaxies

Clicker Question

What kind of measurement does not tell us the mass of a cluster of galaxies?

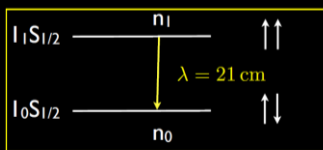
- A. Measure velocities of cluster galaxies
- B. Measure total mass of cluster's stars
- C. Measure temperature of its hot gas
- D. Measure distorted images of background galaxies



The 21-cm Line of Hydrogen

$$\nu_{21cm} = 1,420,405,751.768 \pm 0.001 \text{ Hz}$$

Hyperfine transition of neutral hydrogen

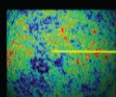


Useful numbers:

- 200 MHz → z = 6
- 100 MHz → z = 13
- 70 MHz → z ≈ 20
- 40 MHz → z ≈ 35
- $t_{\text{Age}}(z = 6) \approx 1 \text{ Gyr}$
- $t_{\text{Age}}(z = 10) \approx 500 \text{ Myr}$
- $t_{\text{Age}}(z = 20) \approx 150 \text{ Myr}$



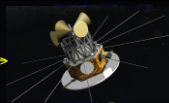
The 21-cm Line in Cosmology



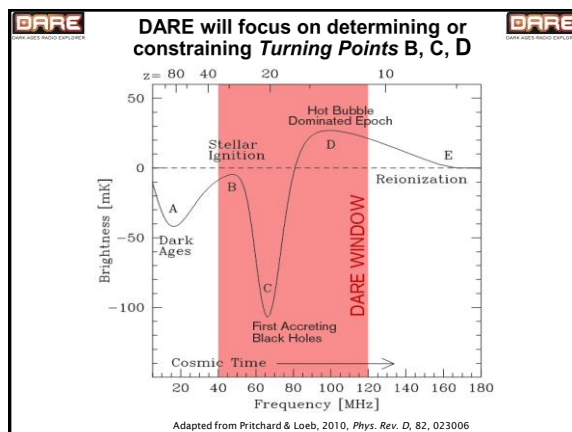
CMB acts as back light



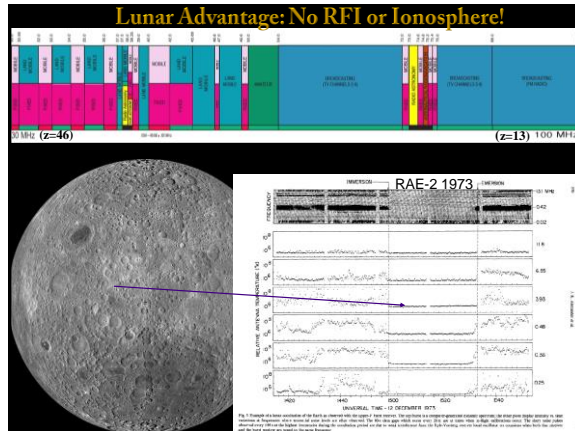
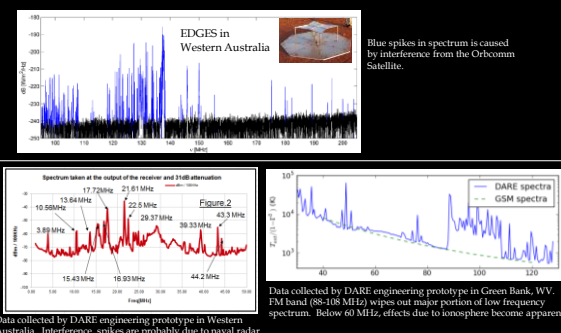
$z = 13$
 $\nu = 1.4 \text{ GHz}$
Neutral gas imprints signal



$z = 0$
 $\nu = 100 \text{ MHz}$
Redshifted signal detected



These observations are best done from the lunar farside which is free of Radio Frequency Interference & the Earth's ionosphere => **DARE will be placed in lunar orbit to access the radio-quiet zone above the lunar farside.**



DARE Spacecraft

Labels: Payload Antenna Support Structure, Payload Dipole Antennas (4) (top view), Payload Antennas Radials (12) (deployable), Spacecraft Bus, 1.8 m² Solar Array, High-gain Antenna, 1.5 meters.

SPACECRAFT CHARACTERISTIC DETAILS

	OSR Mass (kg)	Contingency	Mass w/contingency (kg)
Structures Total	55.5	26%	70.0
Mechanisms Total	0.0		0.0
EPS Total Mass/Power	35.2	32%	46.5
C&DH Total Mass/Power	6.0	5%	6.3
Telecom Total Mass/Power	5.2	9%	6.3
Thermal Total Mass/Power	20.3	25%	25.3
ADC Total Mass/Power	35.9	5%	36.8
Propulsion Total	24.1	12%	26.9
Spacecraft Bus - Dry Total	14.1	20%	25.5
Payload Total Mass/Power	15.0	30%	20.9
Payload - Total Mass	15.0	30%	20.8
Flight System - Dry Total	200.1	21%	241.3
Propellant Total Mass	120.5	0%	120.5
Flight System - Wet Total	320.6	13%	361.8

Burns et al., 2012, *Advances in Space Research*, 49, 433

Launch Option: Secondary Payload on EM-1 Orion MPCV Test Flight

Labels: Launch Abort System Truss Assembly, DARE Spacecraft, Interface Ring, Crew Module Fwd Bay.

DARE easily fits within the Orion payload envelope.

DARE Engineering Prototype: Components

- DARE will operate at low radio frequencies between 40-120 MHz.
- Components: antenna, receiver and spectrometer.
- Instrument Verification Program underway.

Antenna + BALUN (NRAO) → Front End Receiver (NRAO) → Digital Spectrometer (JPL)

1st Generation prototype testing in Green Bank, USA

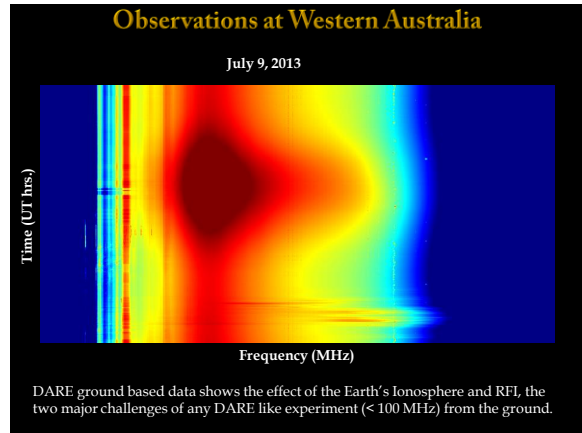
- 2 weeks in March 2012.
- Initial field tests validated the performance of three stages of DARE instrument: antenna, front-end and digital spectrometer.
- 1st generation prototype testing was successful.
- Initial calibrated spectrum showed effects of ionosphere, RFI (FM band).

DARE Prototype in Western Australia

DARE Engineering prototype was deployed on March 21, 2012 at Murchison Radio Observatory (MRO).

DARE Antenna

DARE Instrument Hut



Dark Ages Radio Explorer (DARE)

- DARE is designed to address:
 - When did the First Stars ignite?
 - When did the first accreting Black Holes turn on?
 - When did Reionization begin?
- DARE will accomplish this by:
 - Constructing first sky-averaged spectrum of redshifted 21-cm signal at $11 < z < 35$.
 - Flying spacecraft in lunar orbit & collecting data above lunar farside -- only proven radio-quiet zone in inner solar system.
 - Using high heritage spacecraft & technologies/techniques from DARE Engineering prototype.

