ASTR 1020: Stars & Galaxies October 21, 2013

- Reading: Chapter 19, Sections 19.1-19.2 (for Friday).
- Exam 2 on Wednesday.
- Review Session on Tuesday night at 7 pm, EDUC 155.



Exam 2 will cover

- All material discussed since Exam 1 in class, readings, recitations, and homeworks up through today's class.
- Textbook: Chapters 6, 15, 16, 17, 18.
- MasteringAstronomy Homeworks on The Properties of Stars, Star Birth, and, The Lives of Stars.

The Day of the Exam

- Bring a #2 pencil and eraser or pen.
- One sheet of paper with your notes and study hints.
- Bring a calculator if you think you'll need one.
- Please be prepared to get started right away at 1 pm.



The Stellar Graveyard

Low mass stars → white dwarfs gravity vs. electron degeneracy pressure

High mass stars \rightarrow neutron stars Gravity vs. neutron degeneracy pressure

Even more massive cores → black holes Gravity wins.....

Reading Clicker Question: How do we know pulsars are neutron stars?

- A.The X-ray intensity is too strong to come from a white dwarf.
- B.Pulsars spin too fast to be as large as a white dwarf.
- C.The observed size of the accretion disk is too small to be a white dwarf.
- D.The radiation shows a gravitational redshift that requires a mass greater than the white dwarf limit.

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Today: Neutron Stars

- Gravity vs. Neutron degeneracy pressure
- Size ~ 10 km !!
- =>Crushing gravity at its surface.



Neutron star over New York City























Pulsar "Lighthouses" don't



Clicker Question

Could there be neutron stars that appear as pulsars to other civilizations but not to us?

A. Yes

B. No

Clicker Question

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A. Yes B. No

Neutron Stars in Binary Systems

- Mass transfer: Gravitational potential energy
 → X-ray radiation emission
- X-Ray Binary system, X-ray bursters
- Matter falling through the spinning disk can spin UP the pulsar!



