

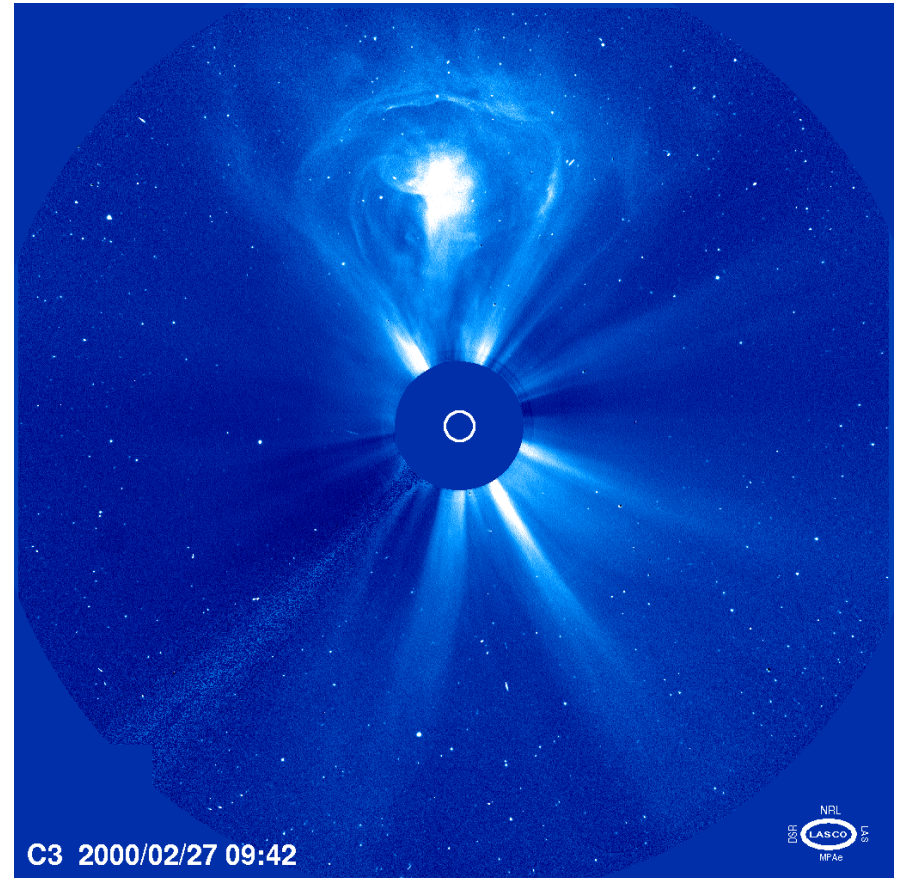
# Heliophysics and Low Frequency Emissions from the Sun

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# Outline

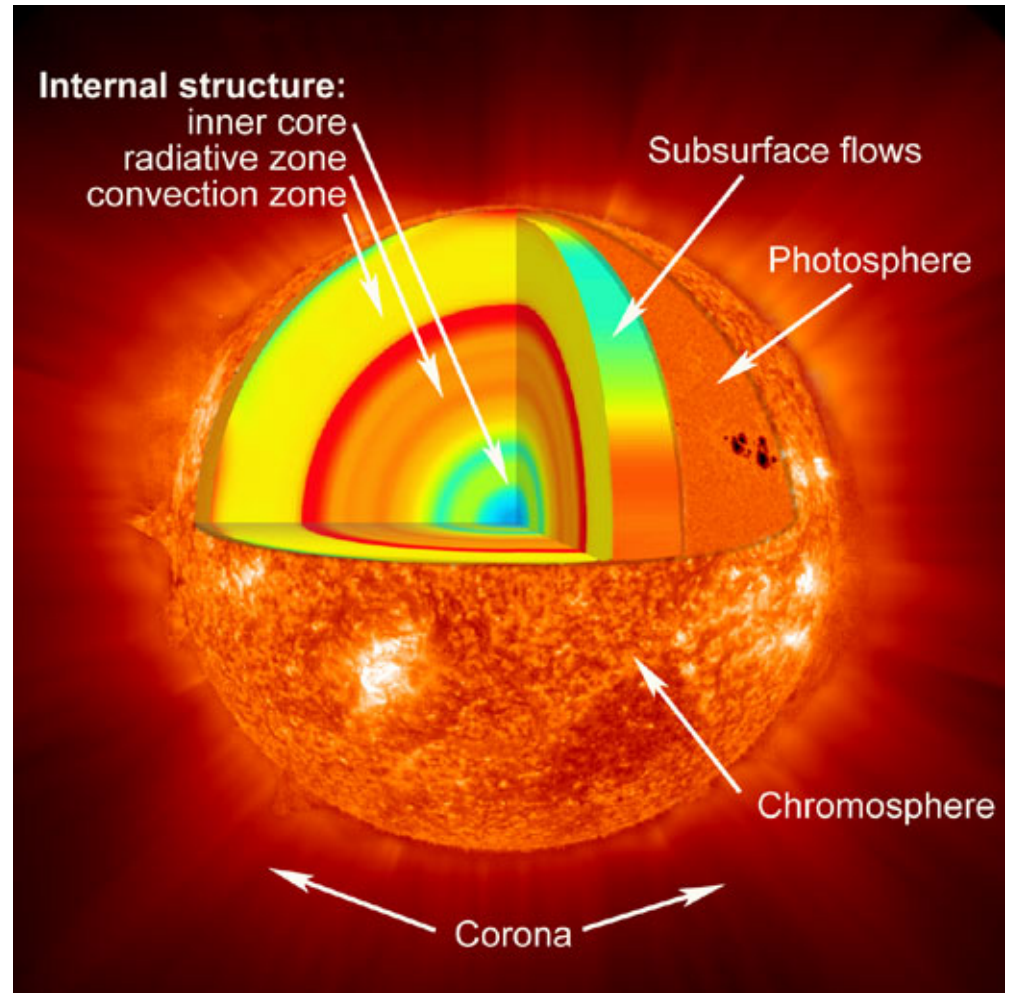
- Review Solar Physics
- Coronal Mass Ejections & other Solar Phenomenon
- Gopalswamy Paper
- Lunar Radio Interferometer





# Our Star - The Sun

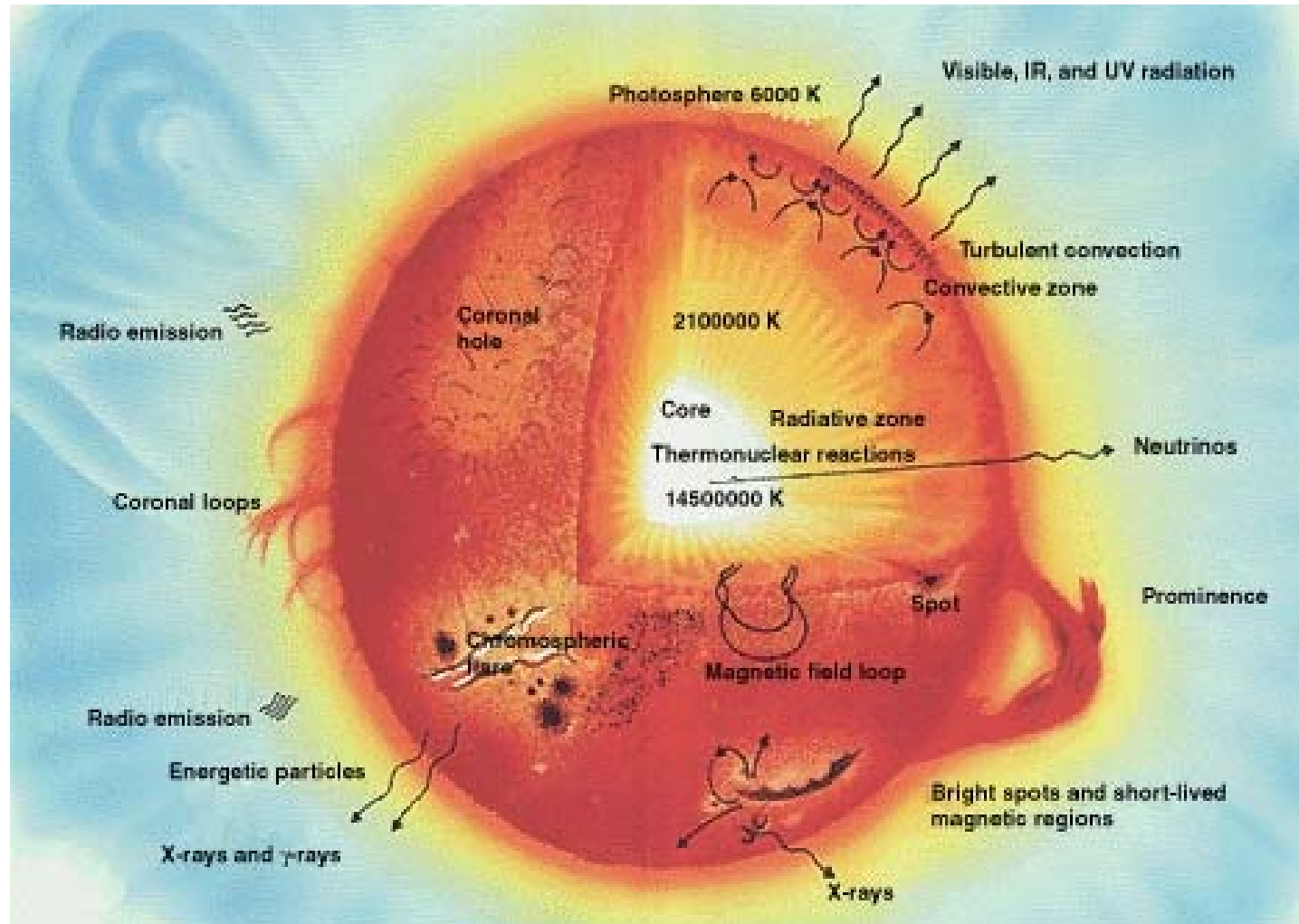
- Not just a ball of hot gas
- Large sphere of plasma
- Differential Rotation
- Complex Layering



# Anatomy of the Sun

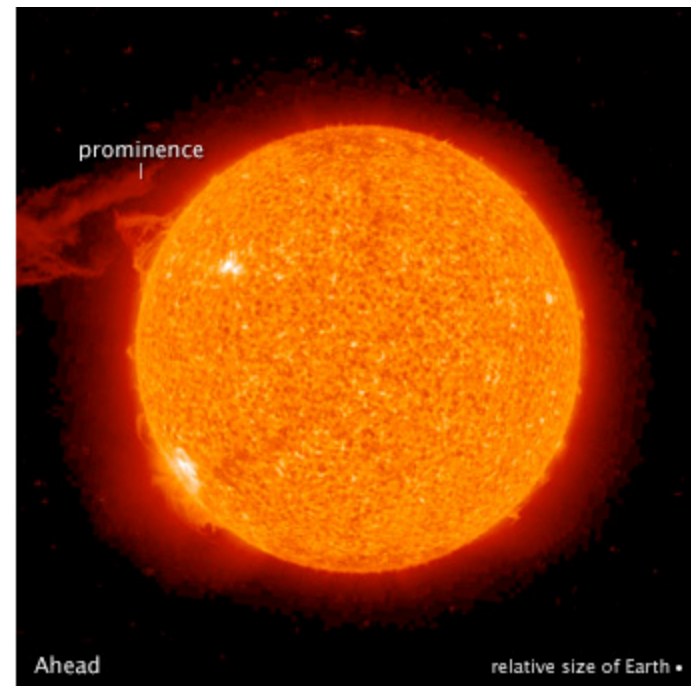
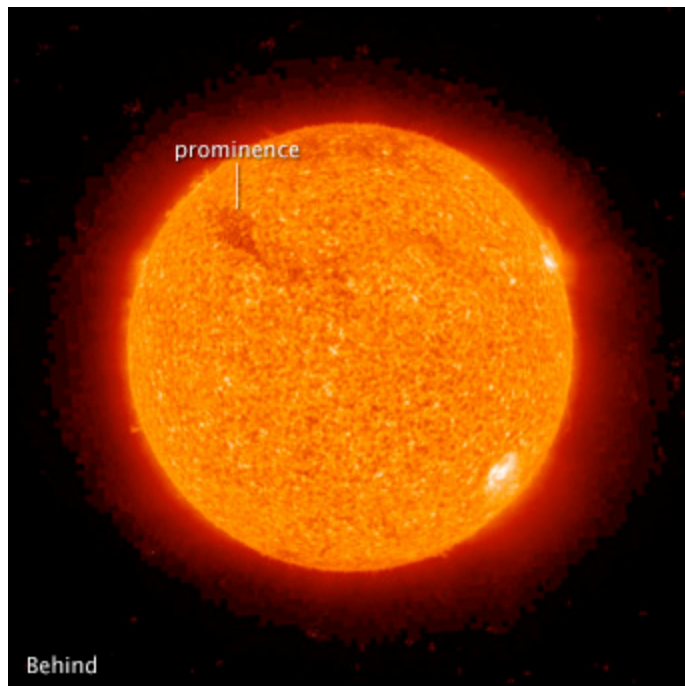
• Region	• Solar Radii	• Temp (Kelvins)
– Core	– .25	– 14 million
– Radiative Zone	– .25 to .7	– 7 to 2 million
– Convective Zone	– .7 to 1	– 5700K
– Photosphere	– Visible Surface	– 6000K
– Atmosphere		
• Chromosphere	• 1 to 1.005	• 4100 to 20,000
• Corona	• ~ 20 (.1AU)	• 1 to 20 million
• Heliosphere	• 50AU - Heliopause	

# Solar Dissection



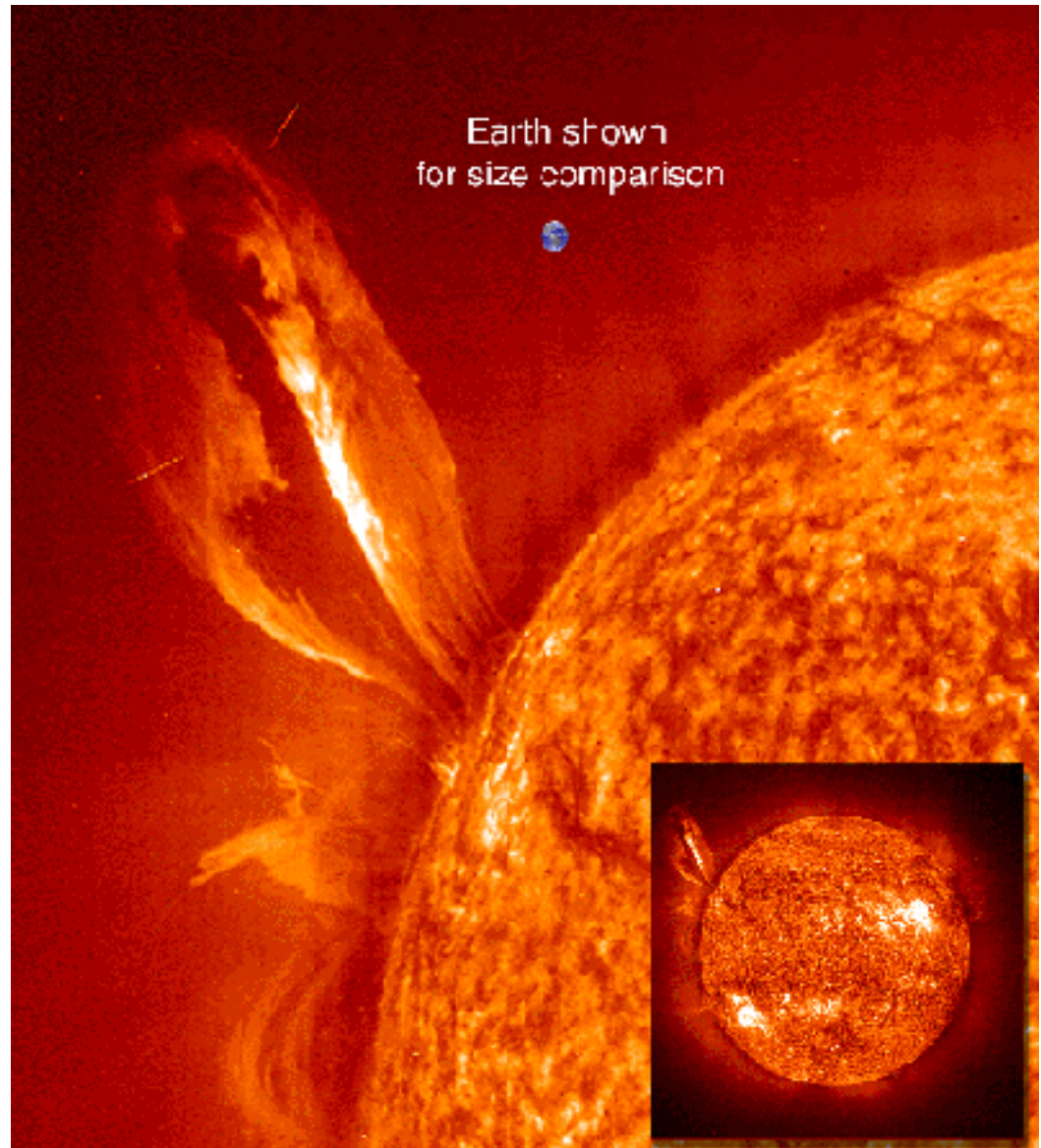
# Solar Prominence

- Cooler temperature plasma
- Material held in loops by magnetic fields
- Solar Filaments are same as Prominence (viewing angle)



# Solar Flares

- Properties
  - Ionized gas large explosions ( $10^{32}$  ergs)
  - Release of magnetic energy
  - Accelerate ions & electrons
  - Emissions across entire spectrum
  - Potential trigger of CME's
  - Vary with solar cycle



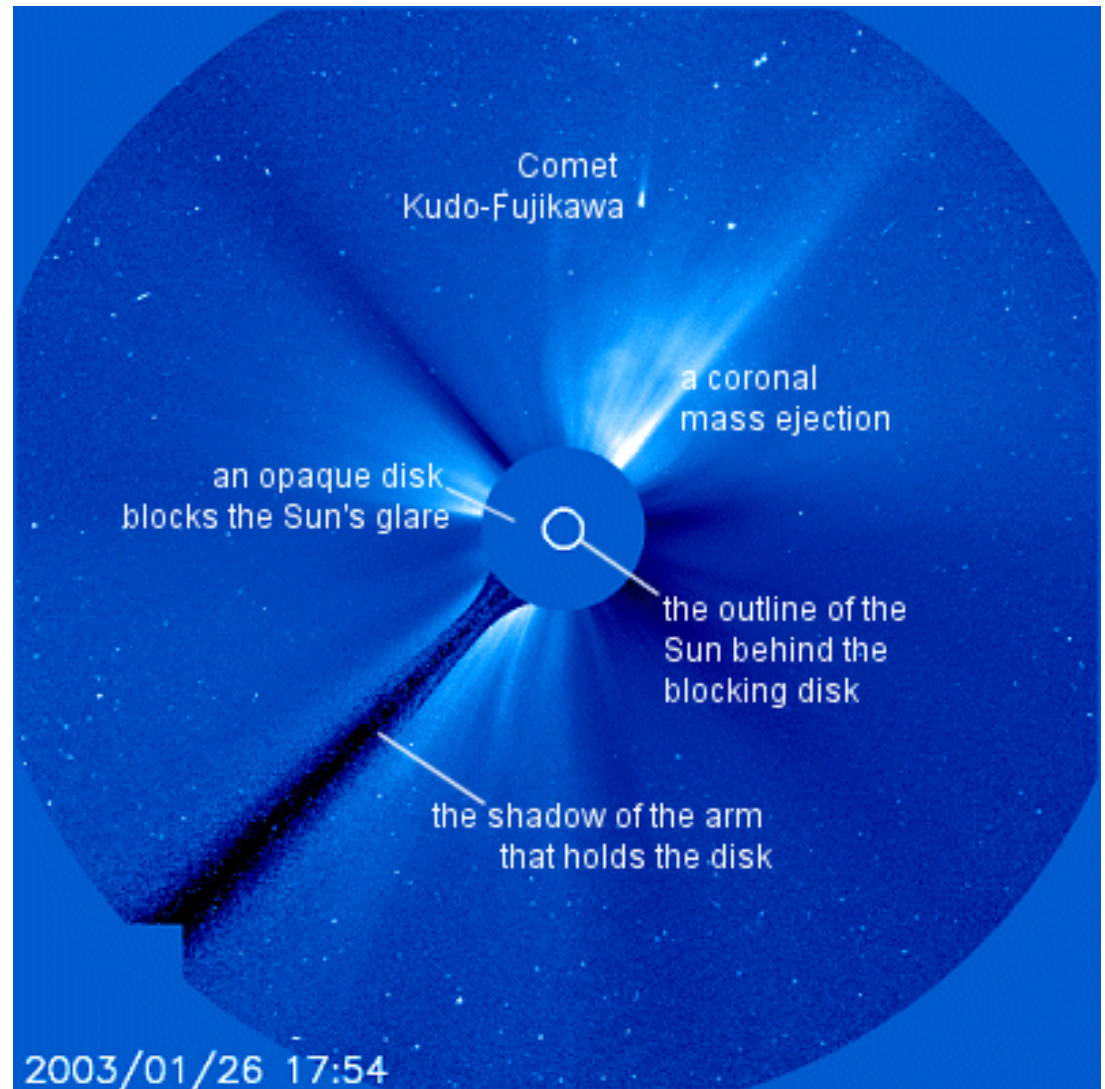


# The Corona

- Invention of Coronagraph
- Why is it so hot?
  - Possible wave heating
    - Dissipating MHD, gravity & sound waves
  - Possible magnetic heating
    - Magnetic reconnection
    - Magnetic Energy

$$\varepsilon = \frac{B^2}{8\pi}$$

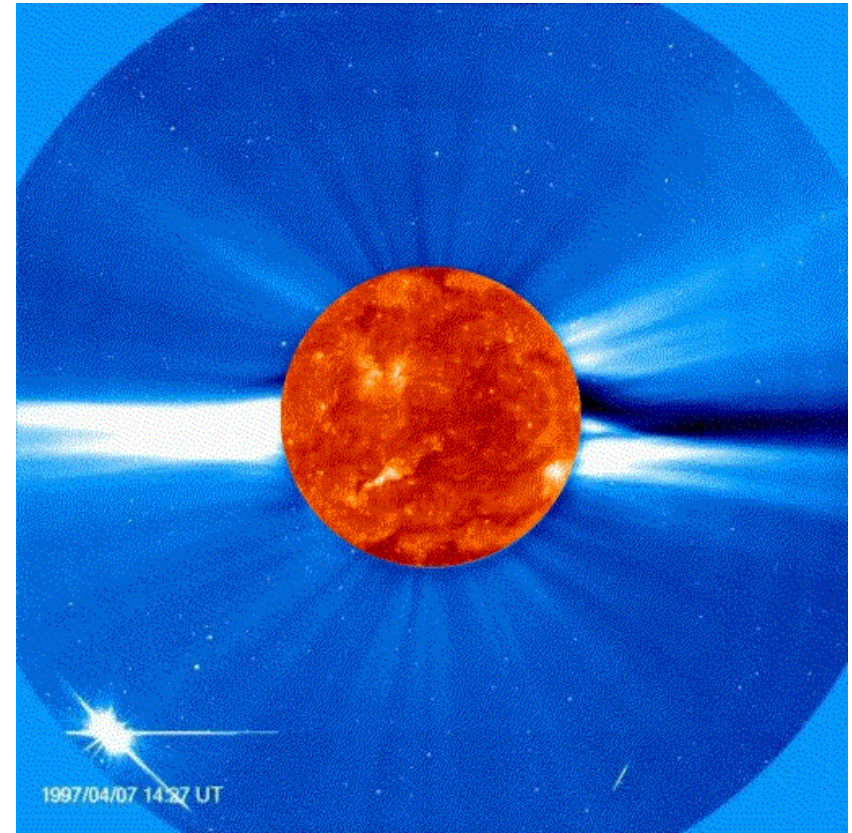
$10^{34}$  ergs



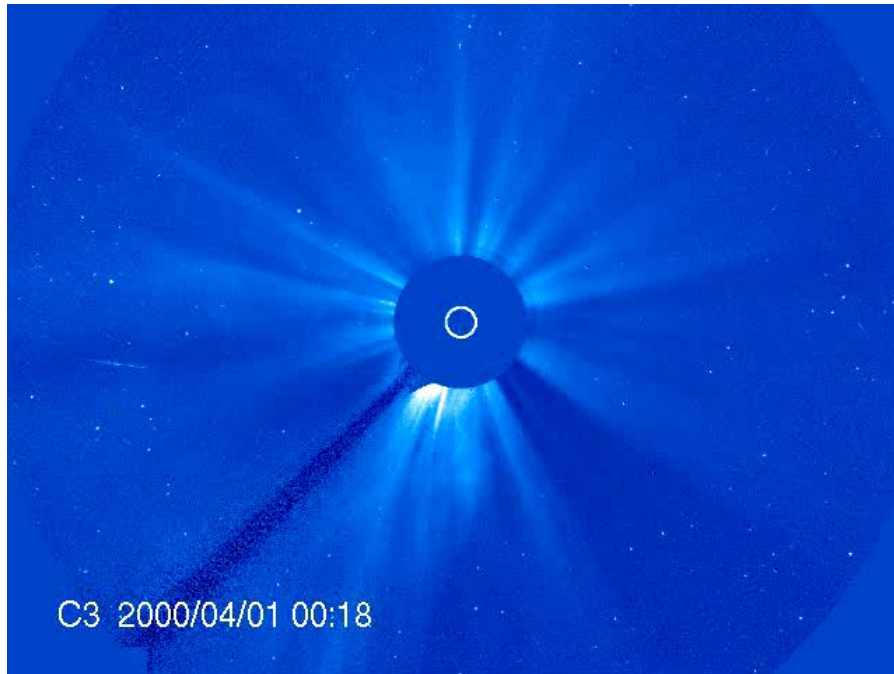


# Coronal Mass Ejections

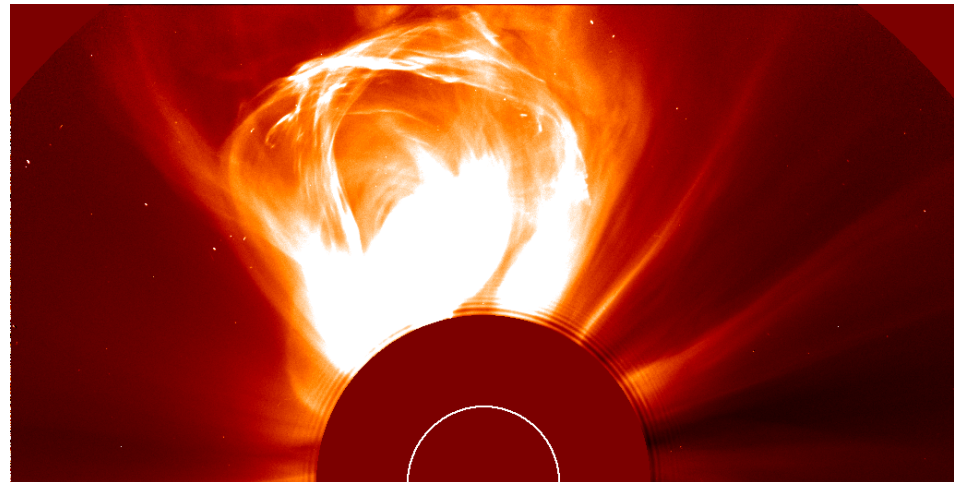
- Gas blown from Corona
  - $10^{15}$  grams of gas (lower limit average)
- Velocity range
  - 20km/s to 3000km/s
- Frequency Occurrence
  - 1/week @ Solar min
  - 2-3/day @ Solar max
- Location
  - Focused on equator during solar min
  - Varying latitudes during solar max
- Origin
  - Correlation to solar flares, prominences & sunspot regions
  - Also occur in absence of the above



# Anatomy of Coronal Mass Ejections



- Explosion from Corona
- Shape of a Croissant
  - Outer Leading Edge
  - Cavity
  - Core



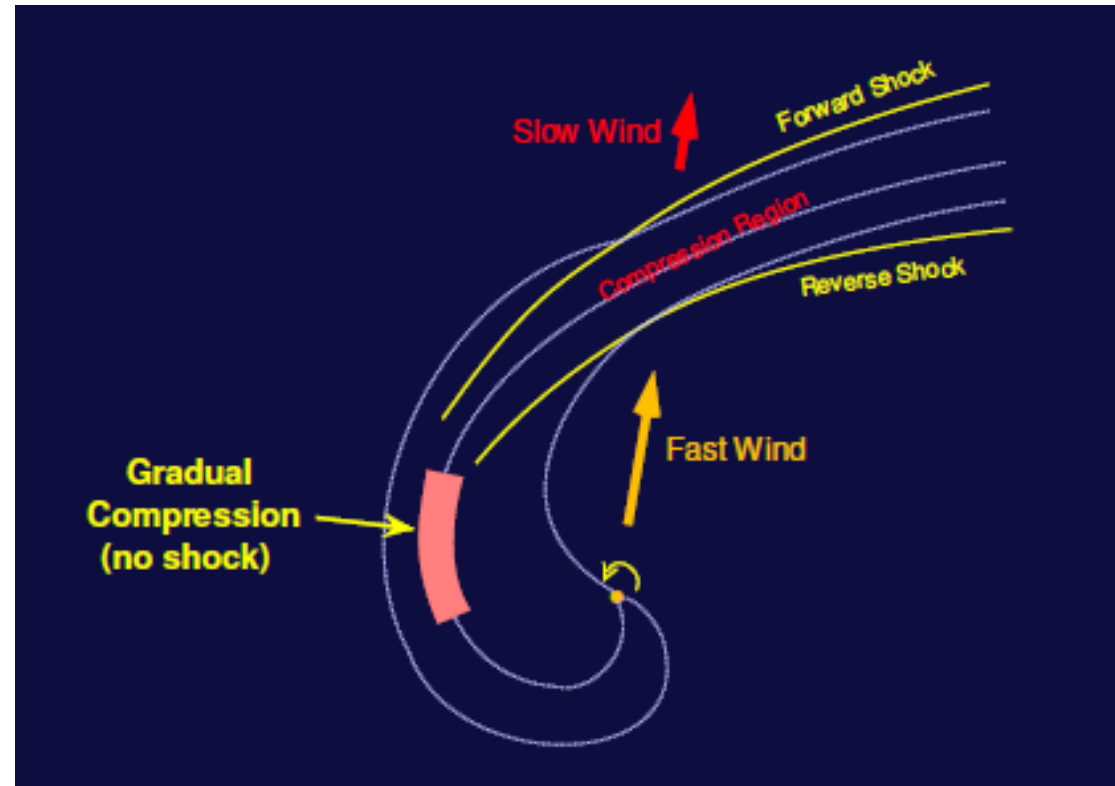
# CME Propagation

- Ejected material reaches millions of Kelvins
- Slow to fast acceleration, then constant velocity
- Velocity can be greater than sound speed of solar wind
  - Solar wind velocity (avg) – 145km/s
- As these velocities increase the mach number of the CME increases
- CME's produce shock fronts



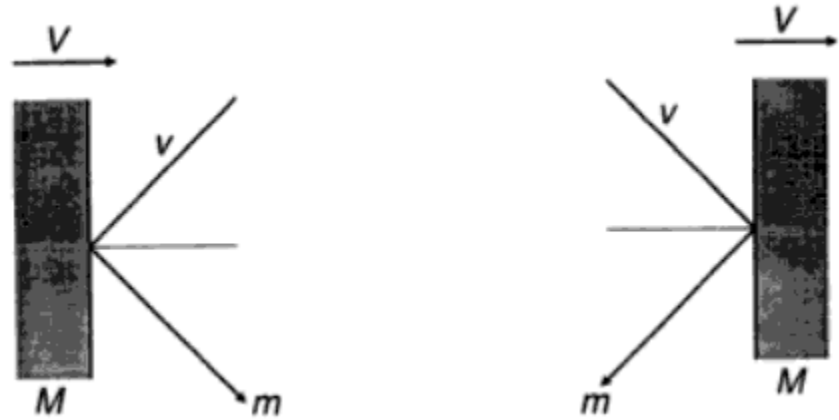
# Fermi Acceleration

- a.k.a. *Diffusive Shock Acceleration*
- Charged particles (ie: electrons) get accelerated by shock front moving through plasma
- Cyclotron or synchrotron radiation occurs



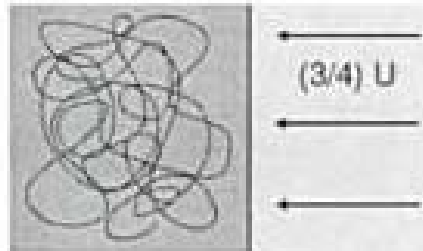
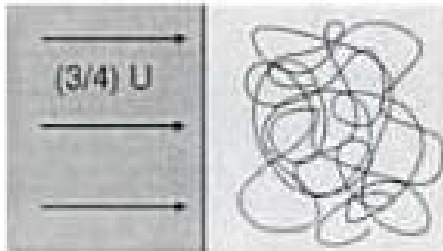
# Fermi Acceleration Mechanism

- Changing reference frames for relativistic particles
- Change of energy increases each occurrence
- Symmetry of head on or following collisions across shock front
- Each boundary crossing increases energy



$$E' = \gamma * (E + Vp \cos \theta)$$

$$E'' - E = \Delta E = \frac{2Vv \cos \theta}{c^2} + 2 * \frac{V^2}{c^2}$$



$$\frac{\partial N(\epsilon)}{d(\epsilon)} \propto \epsilon^{-\eta}$$

# Cyclotron & Synchrotron Radiation

- Photons created from charged particles being deflected by magnetic fields
- Lorentz Force acts on velocity perpendicular to magnetic field and causes emission of radiation and spiraling of particle
- Origin of radio emissions in astrophysics

$$\frac{-dE}{dt} = \frac{\sigma_t B^2 V^2}{c\mu_o}$$



$$\frac{\partial N(\varepsilon)}{d(\varepsilon)} \propto \varepsilon^{-\eta}$$

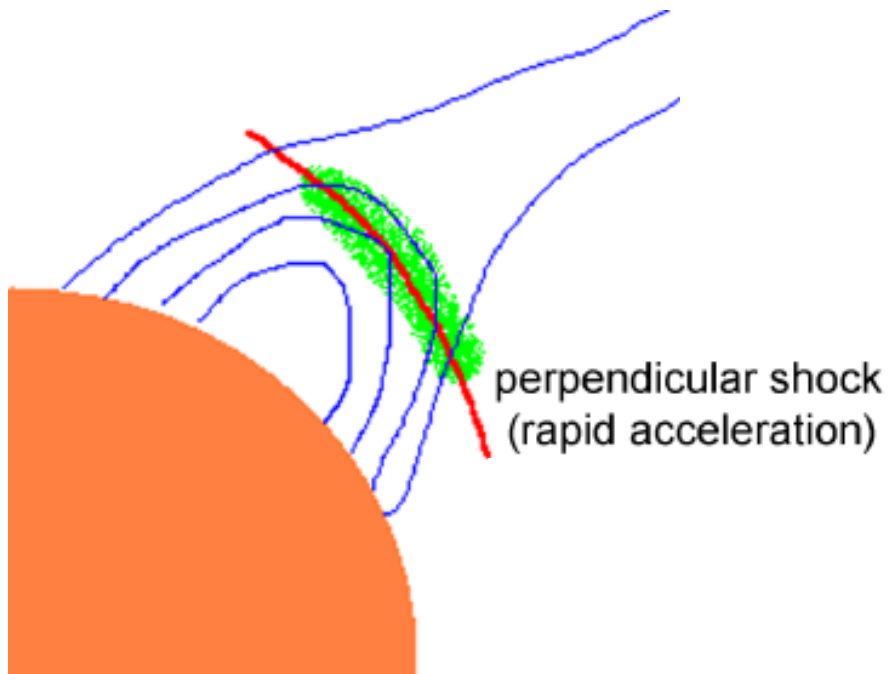


$$\frac{\partial P(\nu)}{d(\nu)} \propto \nu^{-\alpha}$$



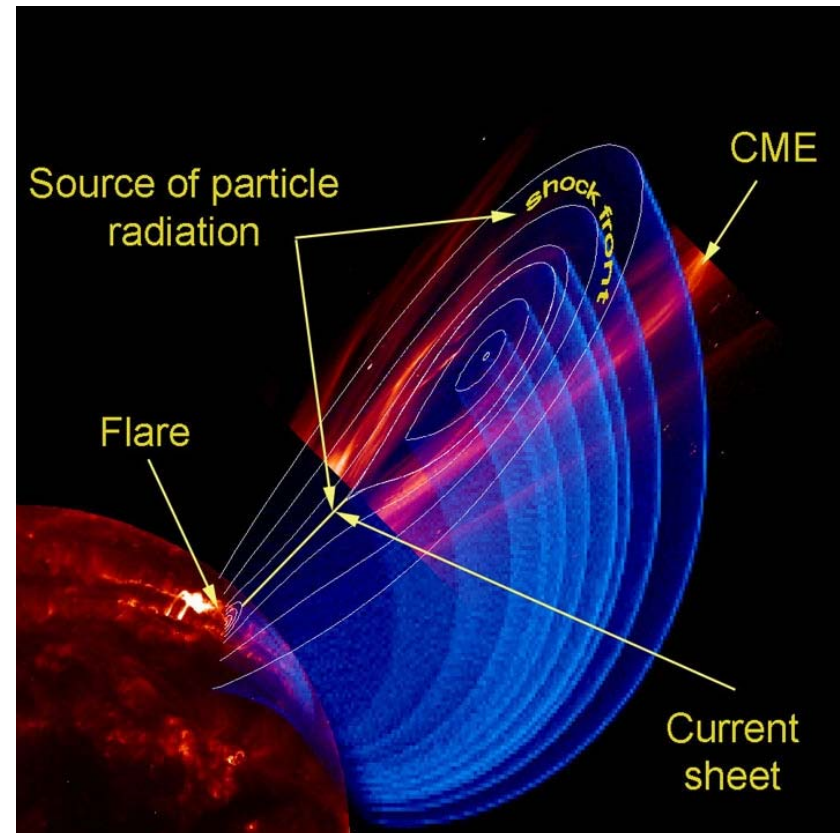
# Shock Acceleration Energies

- Field Orientation
- High energy class
  - MeV to GeV
- Highest Energies
  - 80% speed of light
- SEP event origins?



# Solar Energetic Particles

- Solar Energetic Particles (SEP's)
  - Protons, electrons and ions
  - Can reach 80% speed of light
  - Energy Range keV to GeV
- Origins
  - Solar Flares
  - Shock Fronts in CME's
- Harmful to us and technology above 40MeV
- Need to predict and forecast



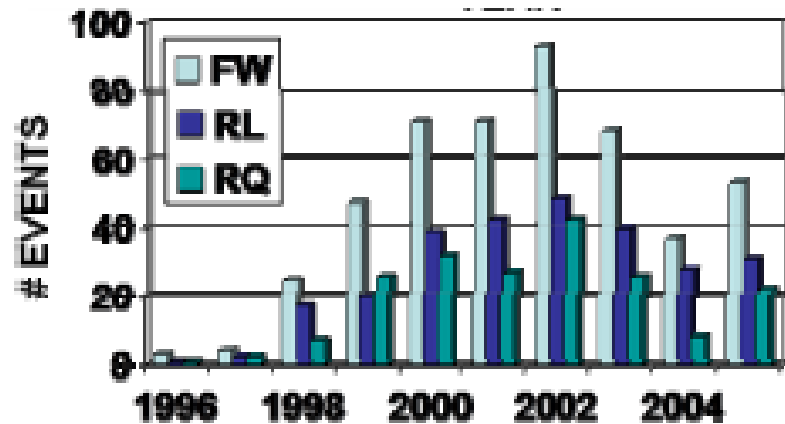
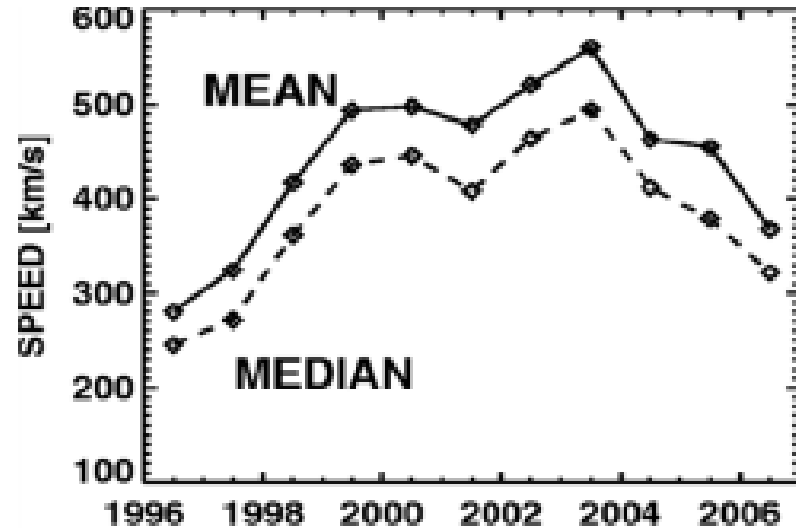
# Paper Intro

- Categorize CME's by velocity and Radio volume (quiet RQ or loud RL)
- Type II Bursts by wavelength
- CME's with Type II bursts are more energetic
- Type II Bursts at longer wavelengths are good indicators of SEP events
- Correlation between CME's and SEP events remains unclear



# Solar Cycle & CME's

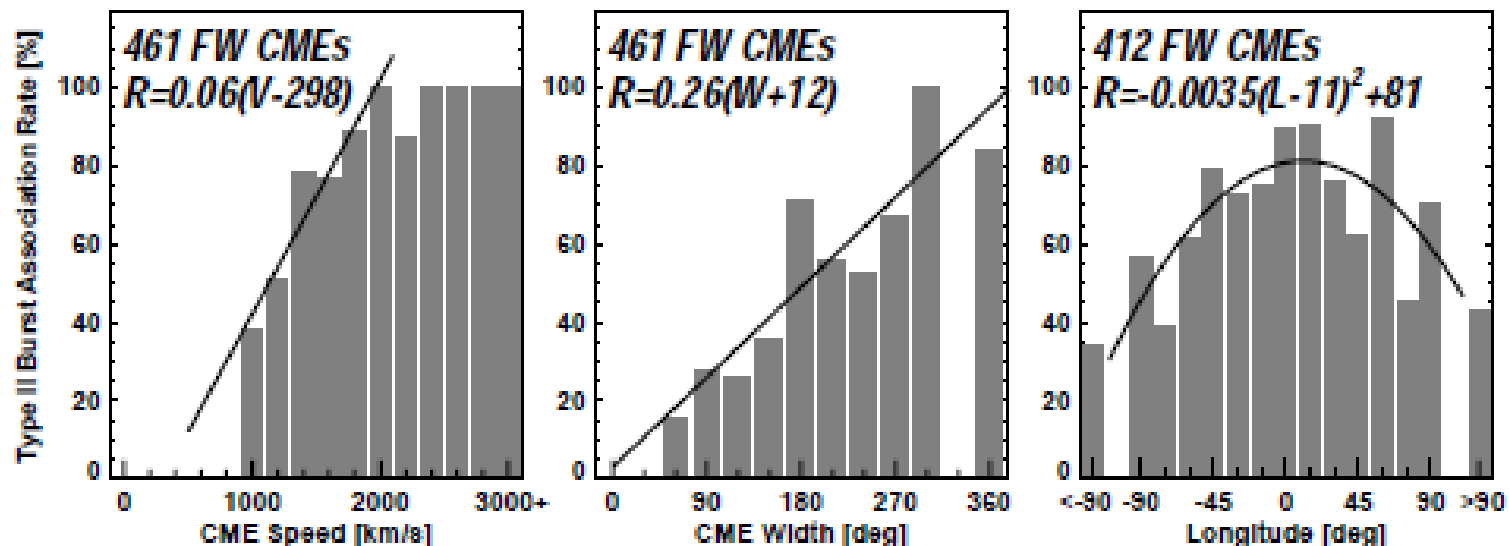
- Solar cycle maximum was 2001-2002
- Velocity and frequency match solar cycle output
- 1999 RQ > RL



# CME's & Type II Bursts

Property	Radio-quiet	Radio-loud
Number of FW CMEs	193 (42%)	268 (58%)
Average speed	1117 km/s	1438 km/s
Average width	86°	89°
Fraction of halos	16%	60%
Median flare size	C6.9	M3.9
Fraction of backside CMEs	55%	25%
East-west asymmetry	-0.02	0.2
Center-to-limb variation	increase	decrease
SEP association <sup>b</sup>	none	55%

- Good correlation seems to exist
- Statistically this relationship is proven
- However not all Type II Bursts are SEP event

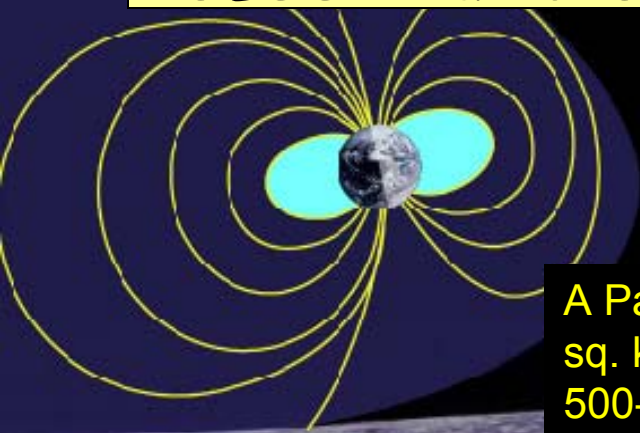


# Paper Findings

- CME's and Radio Emissions are linked
- Why are some CME's radio quiet?
- Why don't all CME's produce SEP events?
  - 1.) They don't produce shocks
  - 2.) They do produce shocks but are 'too' radio quiet to detect
- Statistical analysis of event correlation only produced similar results to Sheeley (1984)



# ROLSS: Radio Observatory for Lunar Sortie Science



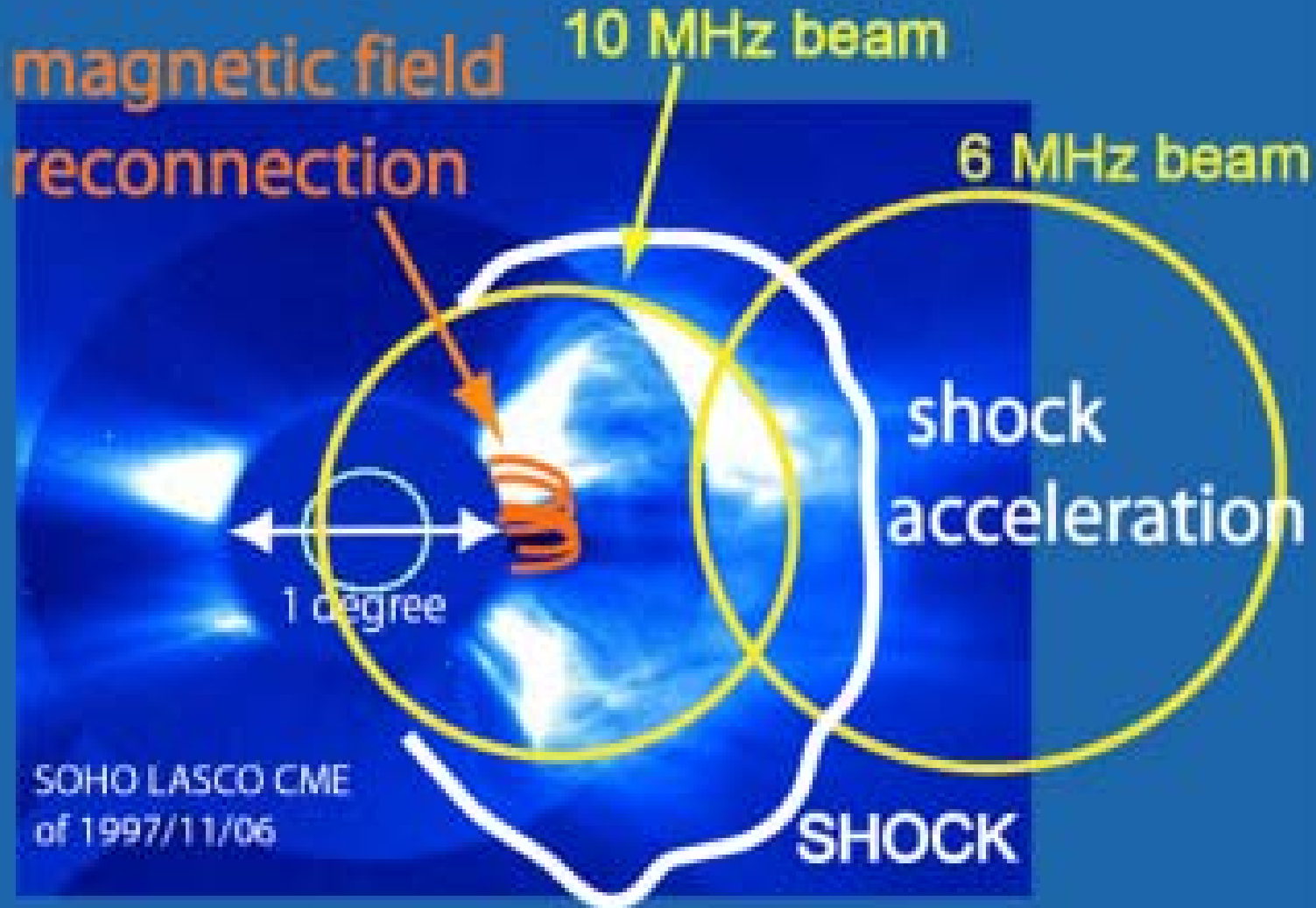
A Pathfinder for a future long-wavelength farside lunar array (10-100 sq. km). Operating at 1-10 MHz (30-300 m). Array consists of three 500-m long arms forming a Y; each arm has 16 antennas.

- Arms are thin polyimide film on which antennas & transmission lines are deposited.
- Arms are stored as 25-cm diameter x 1-m wide rolls (0.025 mm thickness).



Lead Scientists: J. Kasper (CfA) and R. MacDowall (GSFC)

## Complex type III burst source



# Lunar Interferometer & the Sun

- Allow first low frequency imaging of sun
- Locate origin of radio emissions with resolution
- Large collecting area can yield sensitivity to detect weaker radio emissions
- Test current models of CME's and Type II Bursts
- Allow the prediction of Solar Energetic Particle events (SEP's)



# References

- Gopalswamy et al – Coronal mass ejections, type II radio bursts and solar energetic particle events in the SOHO era, Annales Geophysicae; Oct 2008
- Longhair, M.S. – High Energy Astrophysics, Stars, the Galaxy and Interstellar Medium, Vol 2, 1994
- SOHO NASA movie site  
[http://sohowww.nascom.nasa.gov/bestofsoho/Movies/movie\\_s2.html](http://sohowww.nascom.nasa.gov/bestofsoho/Movies/movie_s2.html)
- NRAO website
- Solar Science NASA website

# Discussion

