



# A LUNAR LASER RANGING RETRO-REFLECTOR ARRAY for the 21<sup>st</sup> CENTURY

**Professor Douglas Currie**

University of Maryland, College Park, MD, USA  
NASA Lunar Science Institute, Moffett Field, CA  
INFN – LNF, Laboratori Nazionali di Frascati, Italy



**& The LLRRA-21 Teams**

**Lunar University Network for Astrophysics Research**  
**LUNAR**  
Jack Burns, Director

**Logos on the left side:**

- Georgia Tech (GT)
- MIT Massachusetts Institute of Technology
- Harvard University (VE RI TAS)
- UNM (University of New Mexico)
- UCLA
- UCSD (University of California, San Diego)

**Logos on the right side:**

- NRAO (National Radio Astronomy Observatory)
- GSFC (Goddard Space Flight Center)
- CfA (Harvard-Smithsonian Center for Astrophysics)
- JPL (Jet Propulsion Laboratory)
- Naval Research Laboratory (WASHINGTON, DC)
- University of Maryland (bottom right)

LUNAR Workshop  
October 2010

# LLRRA-21 Teams

## • LSSO Team – NASA

- Douglas Currie Principal Investigator
  - University of Maryland, College Park, College Park m MD, USA
  - NLSI, Moffett Field, CA, USA &
  - INFN-LNF Frascati, Italy
- Bradford Behr
  - University of Maryland, College Park, MD, USA
- Tom Murphy
  - University of California at San Diego, San Diego, CA , USA
- Simone Dell’Agnello
  - INFN/LNF Frascati, Italy
- Giovanni Delle Monache
  - INFN/LNF Frascati, Italy
- W. David Carrier
  - Lunar Geotechnical Institute, Lakeland, FL, USA
- Roberto Vittori
  - Italian Air Force, ESA Astronaut Corps
- Ken Nordtveldt
  - Northwest Analysis, Bozeman, MT, USA
- Gia Dvali
  - New York University, New York, NY and CERN, Geneva, CH
- David Rubincam
  - GSFC/NASA, Greenbelt, MD, USA
- Arsen Hajjan
  - University of Waterloo, ON, Canada

## • INFN-LNF Frascati Team

- Simone Dell’Agnello PI INFN-LNF, Frascati, Italy
- Giovanni Delle Monache INFN-LNF, Frascati, Italy
- Douglas Currie U. of Maryland, College Park, MD, USA
- NLSI, Moffett Field, CA, USA & INFN-LNF, Frascati, Italy
- Italian Air Force & ESA Astronaut Corps
- Claudio Cantone INFN-LNF, Frascati, Italy
- Marco Garattini INFN-LNF, Frascati, Italy
- Alessandro Boni INFN-LNF, Frascati, Italy
- Manuele Martini INFN-LNF, Frascati, Italy
- Nicola Intaglietta INFN-LNF, Frascati, Italy
- Caterina Lops INFN-LNF, Frascati, Italy
- Riccardo March CNR-IAC & INFN-LNF, Rome, Italy
- Roberto Tauraso U. of Rome Tor Vergata & INFN-LNF
- Giovanni Bellettini U. of Rome Tor Vergata & INFN-LNF
- Mauro Maiello INFN-LNF, Frascati, Italy
- Simone Berardi INFN-LNF, Frascati, Italy
- Luca Porcelli INFN-LNF, Frascati, Italy
- Giuseppe Bianco ASI Centro di Geodesia Spaziale “G. Colombo”, Matera,

# What have the Apollo Arrays Done

- The Earth-Moon System Provides an Ideal System
  - To Evaluate Relativity and Einstein's Theory
  - To Understand the Interior of the Moon
- Moon is Massive enough to Resist Drag/Pressure
- Moon is Far Enough to be in a Solar Orbit (Weakly Bound)
- LLR Currently Provides our Tests of:
  - The Weak Equivalence Principle (WEP)\*:  $\Delta a / a < 1.3 \times 10^{-13}$
  - The Strong Equivalence Principle (SEP):  $\eta = 4\beta - \gamma - 1 < 4 \times 10^{-4}$
  - Time-Rate-of-Change of  $G$  to  $< 7 \times 10^{-13}$  per year
  - Inverse Square Law to  $3 \times 10^{-11}$  at  $10^8$  m scales
  - Geodetic Precession to 0.6 %
  - Gravitomagnetism to 0.1 %
  - Initial Definition of Liquid Lunar Core
  - Love Numbers of the Crust
  - Free Librations and  $Q$  of the Moon

# LLRRA-21 Motivations

- Astrophysical Science Motivations
  - Fundamental Incompatibility Between
    - Quantum Mechanics and General Relativity
  - Dark Energy may be Aspect of Large-Scale Gravity
    - Dvali Idea Replaces Normal GR with Leaky Gravity
    - Can be Seen in Precession of Lunar Orbit
  - Dark Matter inspires Alternative Gravity Models (MOND)
    - Further Tests of Inverse Square Law could Confirm or Deny
- Lunar Science Motivations
  - Liquid Core – Dimensions, Shape, Rotation
  - Inner Solid Core – Existence, Size, Rotation
  - Rotational Dynamics – Q, External Impacts

# SHORT HISTORY

- Apollo Lunar Laser Ranging Arrays 1969
  - Thermal and Optical Analysis and Testing
  - McDonald LLR Station
- 2006
  - Return to the Moon
  - Could Address Accuracy Limit
- 2007
  - LSSO for 100 mm CCR
  - Lunar Science Sortie Opportunities
- 2009
  - NLSI > LUNAR at University of Colorado



# ASTRO2010 DECADAL SURVEY

## Gravitational and Particle Physics Panel

***Much is unknown about fundamental theory: Modifications of general relativity*** on accessible scales are not ruled out by today's fundamental theories and observations. It makes sense to look for them by testing general relativity as accurately as possible. Cost-effective experiments that increase the precision of measurement of PPN parameters, and test the strong and weak equivalence principles, should be carried out. For example, improvements in **Lunar Laser Ranging** promise to advance this area.

# ASTRO2010 DECADAL SURVEY

## Gravitational and Particle Physics Panel

- The direct detection of gravitomagnetic effects (the Lense-Thirring precession) from Lageos/Grace, Gravity Probe B, and **lunar laser ranging**.
- The lunar laser ranging verification of the strong equivalence principle to  $10^{-4}$ , meaning that the triple graviton vertex is now known to a better accuracy than the triple gluon vertex.
- Limits on the fractional rate of change of the gravitational constant  $G$  ( $< 10^{-12}$ / Limits on the fractional rate of change of the gravitational constant  $G$  ( $< 10^{-12}$ /yr) from **lunar laser ranging**. Atomic experiments limiting time variation of the fine structure constant to  $10^{-16}$ /yr over periods of several years.
- Experiments that are in progress include the Microscope equivalence principle experiment, the **APOLLO lunar laser ranging** observations, and tests of general relativity using torsion balances and atom interferometry.
- Improved strong and weak equivalence principle limits. Better determination of PPN parameters and  $\dot{G}/G$  from next generation **Lunar laser ranging**

# ASTRO2010 DECADAL SURVEY

## Gravitational and Particle Physics Panel

- A new **Lunar Laser Ranging (LLR)** program, if conducted as a low cost robotic mission or an add-on to a manned mission to the Moon, offers a promising and cost-effective way to test general relativity and other theories of gravity (Figure 8.12). So far, LLR has provided the most accurate tests of the weak equivalence principle, the strong equivalence principle and the constancy in time of Newton's gravitational constant. These are tests of the core foundational principles of general relativity. Any detected violation would require a major revision of current theoretical understanding. As of yet, there are no reliable predictions of violations. However, because of their importance, the panel favors pushing the limits on these principles when it can be done at a reasonable cost. The installation of **new LLR retroreflectors** to replace the 40 year old ones might provide such an opportunity. The panel emphasizes again that its opinion that experiments improving the measurements of basic parameters of gravitation theory are justified only if they are of moderate cost. Therefore, it recommends that NASA's existing program of small- and medium-scale astrophysics missions address this science area by considering, through peer review, experiments to test general relativity and other theories of gravity. **The panel notes that a robotic placement of improved reflectors for LLR** is likely to be consistent with the constraints of such a program. It returns to this recommendation below in the context of a recommendation to augment the Explorer program.



# ASTRO2010 DECADAL SURVEY

## Cosmology and Fundamental Physics Panel

These complex spin-induced orbital effects are the consequences of “frame dragging,” a fundamental prediction of Einstein’s theory that has been probed in the Solar System using Gravity Probe B, LAGEOS satellites, and **Lunar laser ranging**, and has been hinted at in observations of accretion onto neutron stars and black holes.

# LIBRATION PROBLEMS

- Why is there a Problem with the Apollo Arrays
  - Libration in Both Axis of 8 degrees
  - Apollo Arrays are Tilted by the Lunar Librations
  - CCR in Corner is Further Away by Several Centimeters
  - Even Short Laser Pulse is Spread
  - Results in a Range Uncertainty by ~2 cm
  - APOLLO Station of Tom Murphy UCSD
    - Thousands of Returns per Normal Point
    - Root N to Get Range to 1 – 2 millimeters
    - Needs Large Telescope
    - Hard to get Daily Coverage

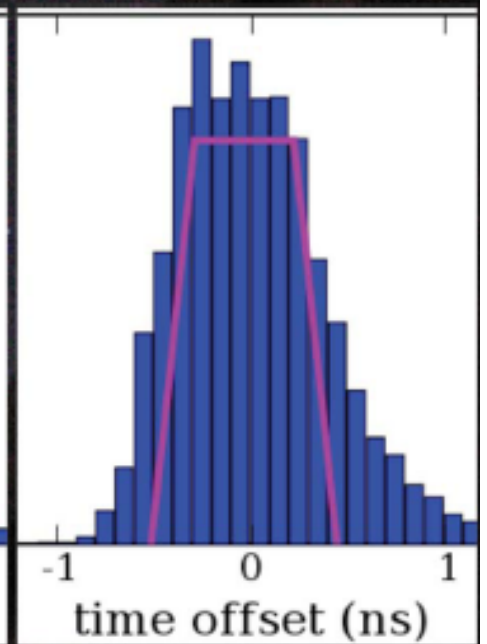
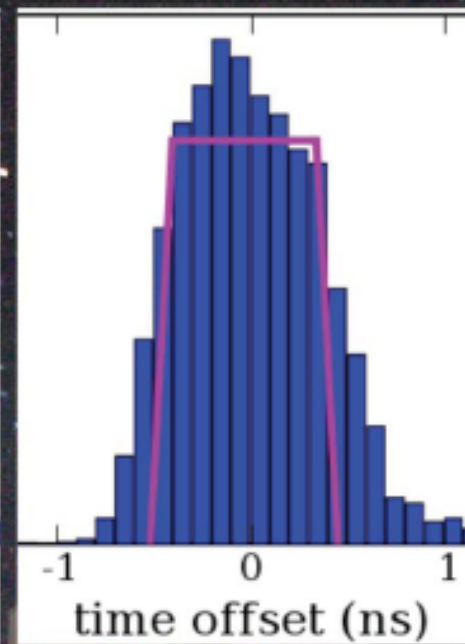
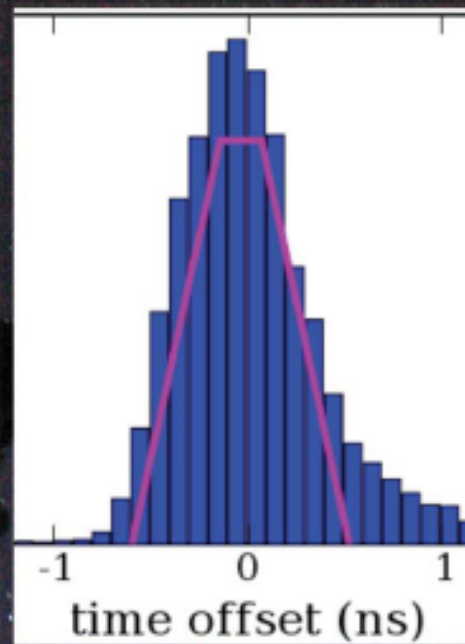
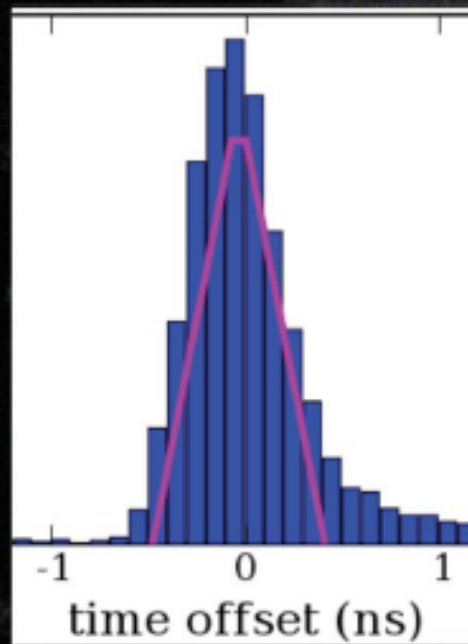
# Sensing Array Size and Orientation

2007.10.28

2007.10.29

2007.11.19

2007.11.20



# LLRRA-2 PROGRAM

- Solid 100 mm Cube Corner Reflector
- 40 Year Heritage, 6.5 TRL
- Program
  - Phase I
    - Surface Emplacement
    - Supports Sub Millimeter Single PhotoElectron Ranging
    - 2012 – 1013
  - Phase II
    - Anchored Emplacement
    - Supports Ranging at less than 100 microns
    - 2016 or Later

# CHALLENGES for SOLID CCR

- Fabrication of the CCR to Required Tolerances
- Sufficient Return for Reasonable Operation
  - Ideal Case for Link Equation
- Thermal Distortion of Optical Performance
  - Absorption of Solar Radiation within the CCR
  - Mount Conductance - Between Housing and CCR Tab
  - Pocket Radiation - IR Heat Exchange with Housing
  - Solar Breakthrough - Due to Failure of TIR
- Stability of Lunar Surface Emplacement
  - Problem of Regolith Heating and Expansion
  - Drilling to Stable Layer for CCR Support
  - Thermal Blanket to Isolate Support
  - Housing Design to Minimize Thermal Expansion

# CCR FABRICATION CHALLENGE

- CCR Fabrication Using SupraSil 1 Completed
- Specifications / Actual
  - Clear Aperture Diameter - 100 mm / 100 mm
  - Mechanical Configuration - Expansion of Our APOLLO
  - Wave Front Error - 0.25 / 0.15 [  $\lambda/6.7$  ]
  - Offset Angles
    - Specification
      - 0.00", 0.00", 0.00" +/-0.20"
    - Fabricated
      - 0.18", 0.15", 0.07"
- Flight Qualified
  - with Certification

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6 October 2010



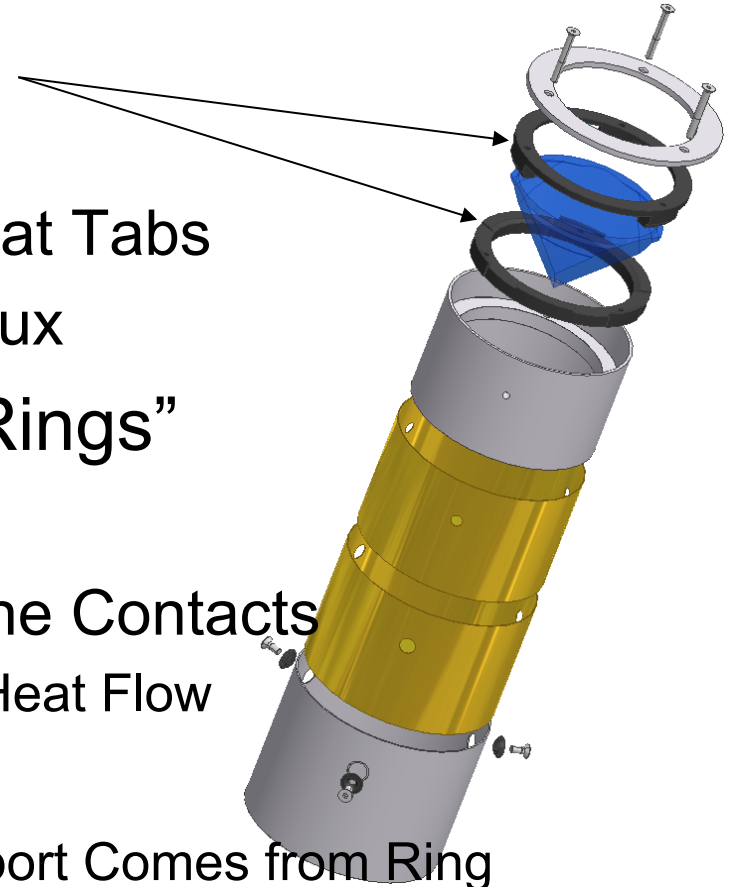
# THERMAL ANALYSIS – THEORETICAL

## Solar Absorption within CCR

- Solar Heat Deposition in Fused Silica
  - Solar Spectrum – AMO-2
  - Absorption Data for SupraSil 1/311
  - Compute Decay Distance for Each Wavelength
  - Compute Heat Deposition at Each Point
    - Beer's Law
  - Thermal Modeling Addresses:
    - Internal Heat Transport and Fluxes
    - Radiation from CCR to Space
    - Radiation Exchange with Internal Pocket Surroundings
    - Mount Conduction into the Support Tabs

# MOUNT CONDUCTANCE

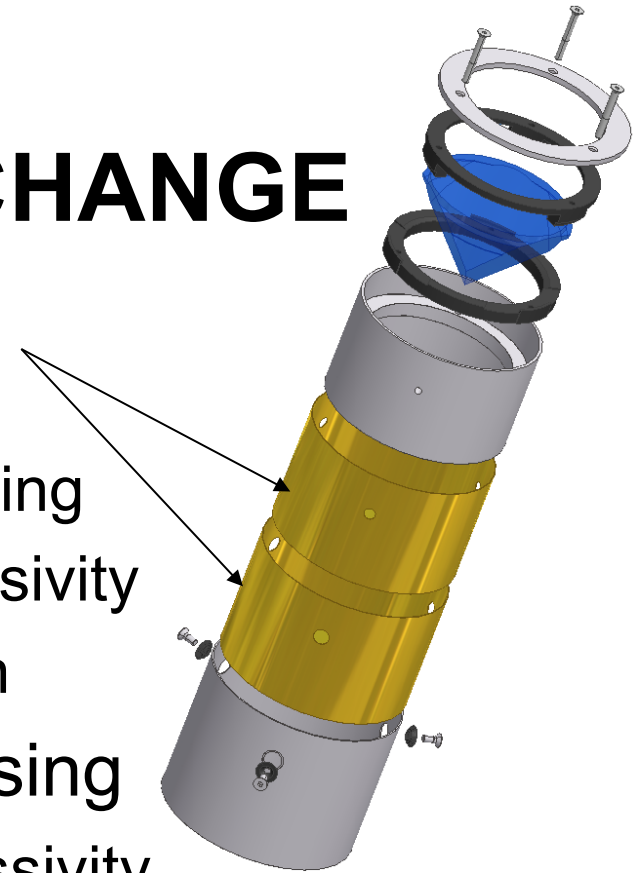
- Challenge:
  - Heat flow from Housing to CCR at Tabs
  - Optical Distortion due to Heat Flux
- Support of CCR with KEL-F “Rings”
  - Intrinsic Low Conductivity
  - Use of Wire Inserts with Only Line Contacts
    - Line Contact of Support Reduces Heat Flow
  - Supports Launch Environment
    - KEL-F Wire Compresses and Support Comes from Ring
- Estimated (to be Validated in SCF) 1 Milli-W/°K



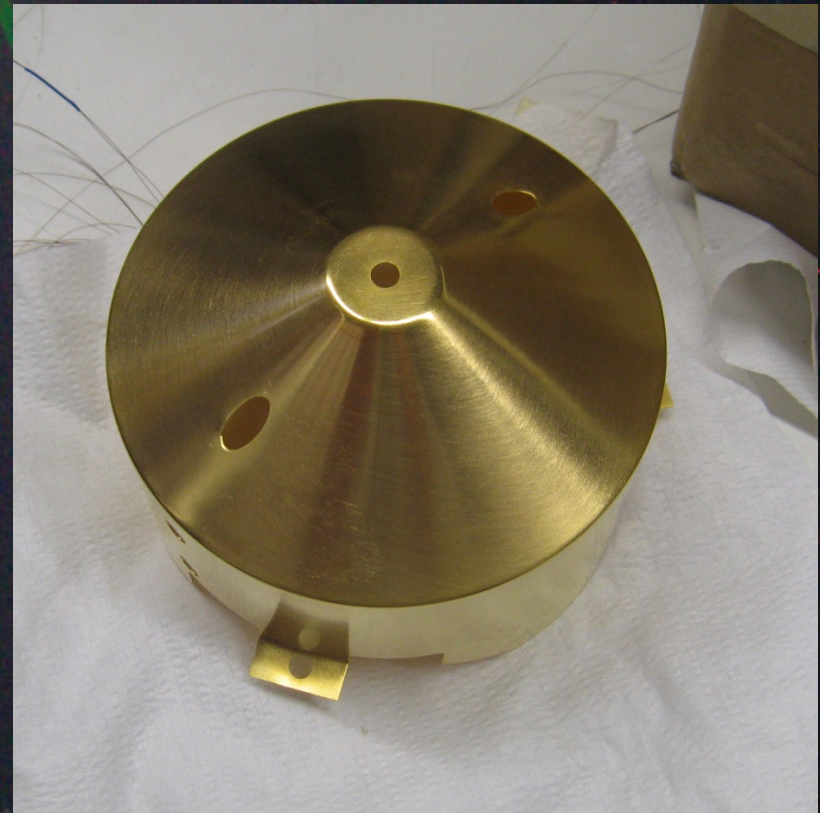
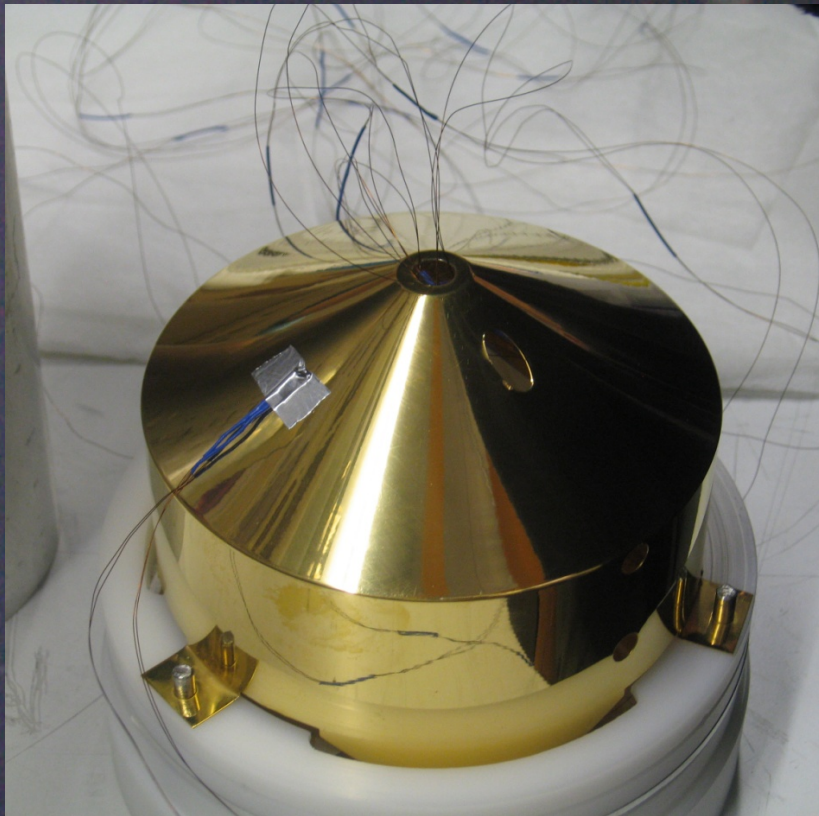


# POCKET RADIATION EXCHANGE

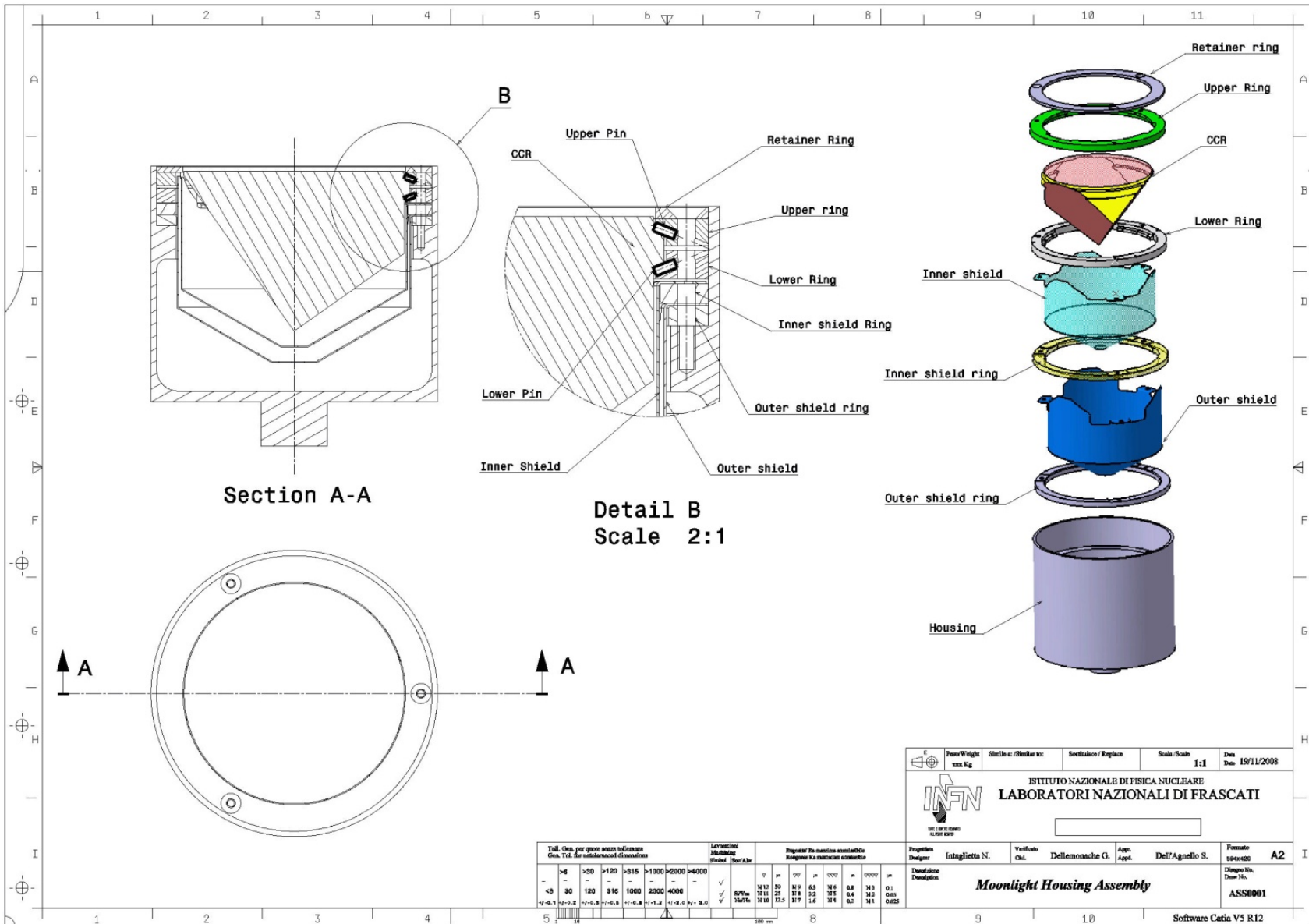
- Challenge:
  - IR Radiation Between CCR & Housing
  - SiO<sub>2</sub> Has High IR Absorptivity/Emissivity
  - Heat Flux Causes Optical Distortion
- Isolation Between CCR and Housing
  - Low Emissivity Coatings – 2% Emissivity
  - Successive Cans or Multiple Layers
- Simulation Indicates Isolation is Effective
- Thermal Vacuum Chamber Validation
  - In April 2009 at SCF at INFN/LNF at Frascati



# INNER & OUTER THERMAL SHIELDS



# LLRRA-21 CONFIGURATION



Part Weight 2224 Kg	Similar to:	Substitute / Replace	Scale / Scale 1:1	Des. Date 19/11/2008
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ISTITUTO NAZIONALE DI FISICA NUCLEARE  
LABORATORI NAZIONALI DI FRASCATI

INGEGNERIA N. Verificato Da: Dellemonache G. Appr. Applt. Dell'Agnetto S. Formato: sbr4x20 A2

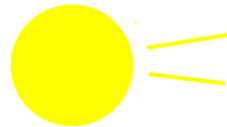
Toll. Dim. per ogni stato di tensione	Dimensioni	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza
Dim. Tol. per tolleranza	Dimensioni	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza	Limite di tolleranza
>0	>120	>315	>1000	>2000	>4000	0.1	0.15	0.2	0.3	0.4	0.5	0.6	0.7	0.8	0.9	1.0	1.2	1.5	2.0	2.5
<0	30	120	315	1000	2000	0.05	0.075	0.1	0.15	0.2	0.3	0.4	0.5	0.6	0.7	0.8	0.9	1.0	1.2	1.5
±0.1	±0.15	±0.2	±0.3	±0.4	±0.5	±0.05	±0.075	±0.1	±0.15	±0.2	±0.3	±0.4	±0.5	±0.6	±0.7	±0.8	±0.9	±1.0	±1.2	±1.5

Programma Disegnato	Intaglietta N.	Verificato Cal.	Dellemonache G.	Appr. Applt.	Dell'Agnetto S.	Formato sbr4x20	A2
Moonlight Housing Assembly						Disegno No. ASS0001	

# ORBITAL THERMAL EVOLUTION

- Simulation Performed with

- Thermal Desktop
  - C&R Technologies
- IDL
- Code V



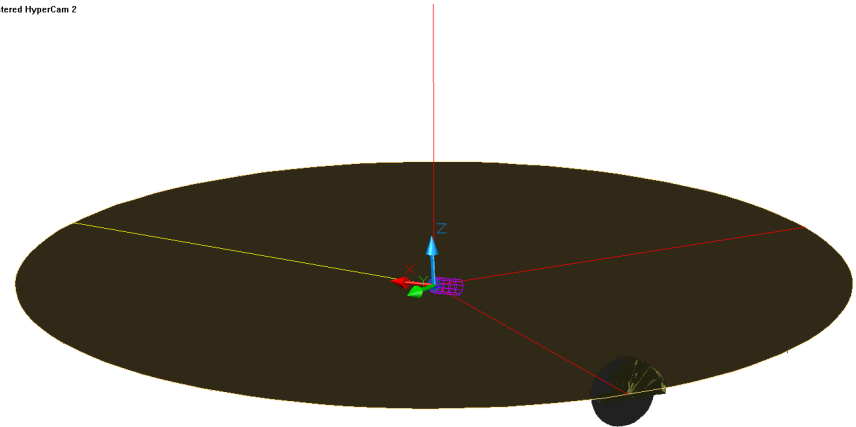
- Initial Analysis

- “Steady State” Behavior
- Fixed Elevation of Sun

- But During Lunar Month

- Changing Illumination
  - Both Intensity and Angle
- For CCR/Housing/Thermal Blanket/Regolith
  - Some Time Constants are Longer than a Month
- Analysis of Behavior of Face to Tip Temperature Difference

Unregistered HyperCam 2

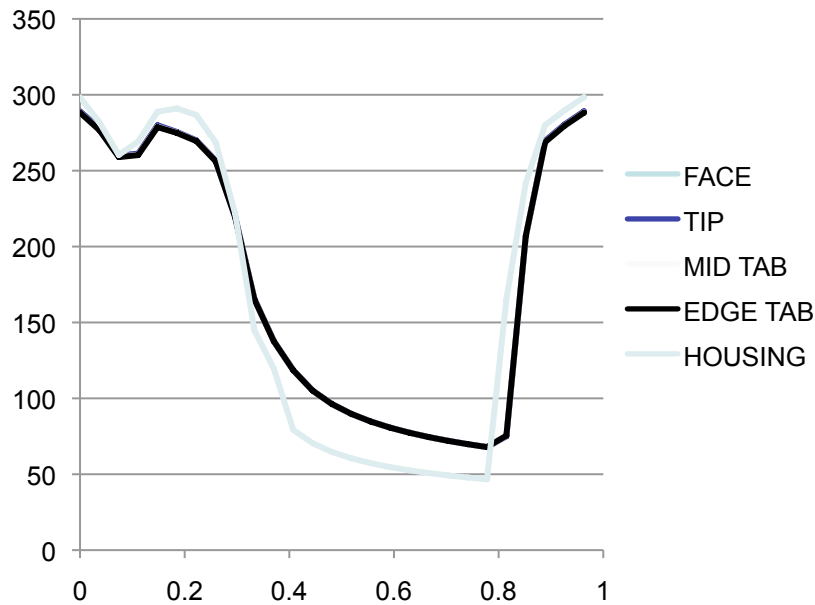


# LUNAR SURFACE EMPLACEMENT

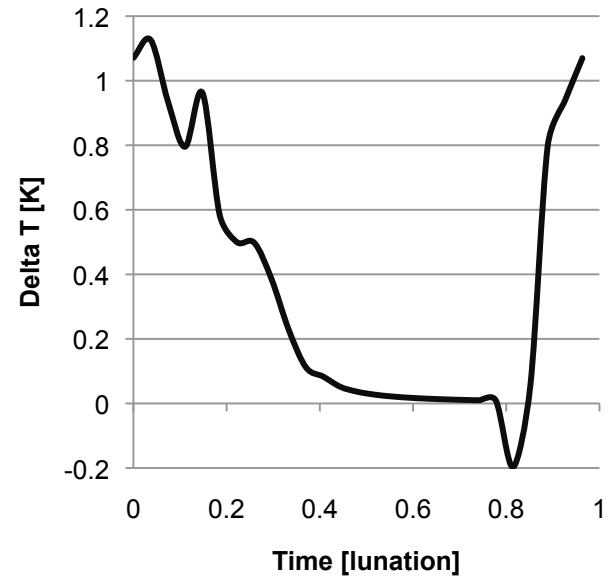
- CCR Optical Performance at Sub-Micron
  - Want to Assure as Much of This as Possible
- We Have Sufficiently Strong Return
- Emplacement Issues - Diurnal Heating of Regolith
  - ~ 400 Microns of Lunar Day/Night Vertical Motion
- Solutions – Dual Approach for Risk Reduction
  - Drill to Stable Layer and Anchor CCR to This Level
    - ~ one meter – Apollo Mission Performed Deeper Drilling
    - ~ 0.03 microns of motion at this depth
  - Stabilize the Temperature Surrounding the CCR
    - Multi Layer Insulation Thermal Blanket – 4 meters diameter
    - Support Rod Sees a Constant Temperature Environment

# TYPICAL THERMAL SIMULATION RESULTS

## ● Temperatures During Lunation

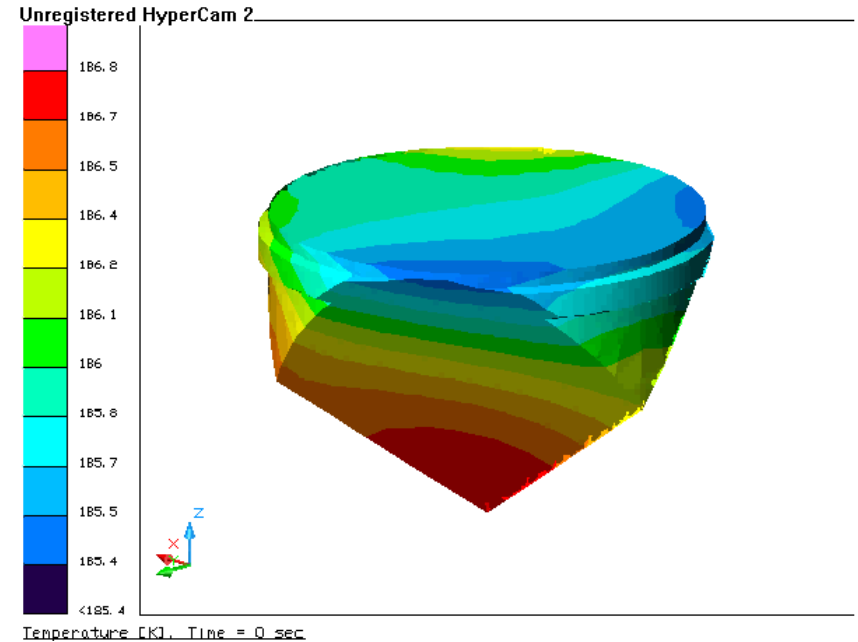
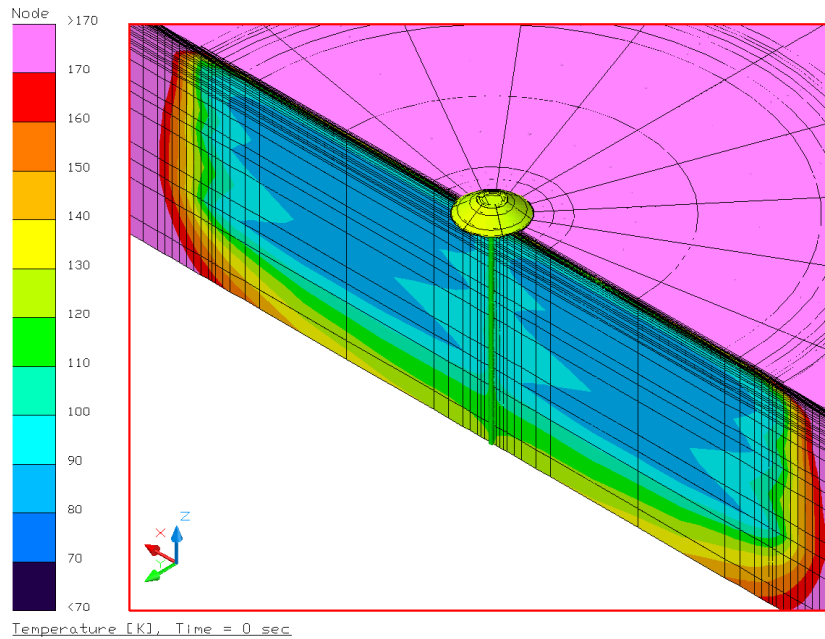


## Tip to Face Temperature Difference



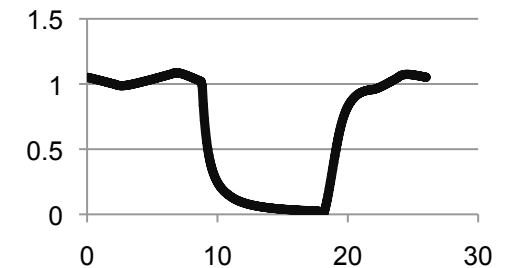
# FULL THERMAL SIMULATION

Unregistered HyperCam 2



Regolith from Apollo HFE,  
Thermal Blanket  
Current Design Housing

Temperature Distribution  
in CCR and Tip to Face  
Temperature Difference



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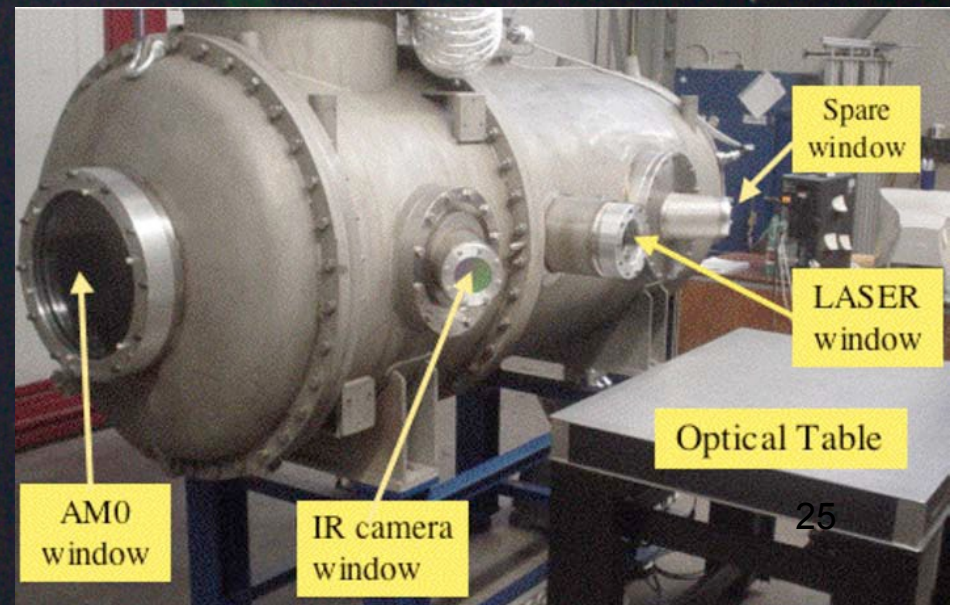
# LLRRA-21 PACKAGE





# CURRENT STATUS

- Preliminary Definition of Overall Package
- Completed Preliminary Simulations
  - LSSO – Lunar Science Surface Opportunities
  - Thermal (CCR, Regolith, Housing), Optical
- Completed Phase I Thermal Vacuum Tests
  - Solar Absorption Effects on CCR
  - CCR Time Constants –
    - IR Camera – Front Face
    - Thermocouples – Volume
    - Preliminary Optical FFDP



# SURFACE DEPLOYMENT

- Issues
  - CCR Should Point Toward Earth “Center”
  - Maintain Clocking Angle to Handle Sun Break-through
  - Handle Longitudinal (toward earth) Tilt of Surface
  - Handle Azimuthal Tilt of Surface
- Requirements
  - Self Orienting Procedure to Keep Clocking Angle
  - Longitudinal (Elevation) Self Orientation
  - Azimuth Angle Adjusted by Arm
    - Calibrated by Goniometer (Sun Dial)

# Apollo 17

## Jack Schmidt & Gene Cernan



# PNEUMATIC PROBOSCIS SYSTEM

Chris Zacny – Honeybee, Inc.




# MISSION OPPORTUNITIES

&

## Acknowledgements

- Possible Roles for 100 mm Solid CCR Retroreflector
  - NASA
    - First Manned Landing - LSSO / NASA Program
    - International Lunar Network (ILN) Anchor Nodes – NLSI /U of Colorado
    - Lunette Discovery Proposal
    - Lunar Express – Lockheed Martin
  - Italian Space Agency & INFN
    - MAGIA – ASI & INFN
      - Proposed ASI Lunar Orbiter to Carry a 100 mm Solid CCR
    - Italian ILN Retroreflector Instrument
      - MoonLIGHT-ILN INFN Experiment – Just Approved



Thank You!

Any Questions?